Probing background with Method of Moments for $B^0 \to K^* \mu \mu$



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- Reminder
- 2 Theory introduction
- Background regions
- 4 Results
- Summary



Plan

Method of moments:

- Last meeting showed how orthogonality of the does magic for method of moments.
- ② Using toy MC (experimental math) checked the errors estimates.
- 3 Checked that it does not suffer from boundary conditions.
- Many thanks to Tom for checking all my calculations.

For today:

• How this method behaves in terms of background?



What do we start with

Let's assume for simplicity we have our pdf:

$$\frac{d^4\Gamma}{\Gamma dq^2 d\cos\theta_k d\cos\theta_l d\phi} = \frac{9}{32\pi} (\frac{3}{4} (1 - F_l) \sin^2\theta_k + F_l \cos^2\theta_k + (\frac{1}{4} (1 - F_l) \sin^2\theta_k + F_l \cos^2\theta_k + (\frac{1}{4} (1 - F_l) \sin^2\theta_k + F_l \cos^2\theta_k + F_l \cos^$$

What did we assume:

- S_{1x} , J_{2x} can be parametrized by F_I .
- $S_{6c} = 0$.
- In short what was in the paper.



Obtained moments 1

Lets see how this works in practice:

$$\frac{d^3\Gamma}{\Gamma d\cos\theta_k d\cos\theta_l d\phi} \sin^2\theta_k = \frac{2}{5} (2 - F_l) \tag{2}$$

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k} dcos\theta_{l} d\phi} \cos^{2}\theta_{k} = \frac{1}{5} (1 + F_{l})$$
(3)

$$\frac{d^3\Gamma}{\Gamma d\cos\theta_k d\cos\theta_l d\phi} \sin^2\theta_k \cos 2\theta_l = -\frac{2}{25} (2 + F_l) \tag{4}$$

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}\cos^{2}\theta_{k}\cos 2\theta_{l} = -\frac{1}{25}(1+8F_{l})$$
 (5)

$$\frac{d^3\Gamma}{\Gamma d\cos\theta_k d\cos\theta_l d\phi} \sin^2\theta_k \cos\theta_l = \frac{2S_{6s}}{5} \tag{6}$$



Obtained moments 2

$$\frac{d^3\Gamma}{\Gamma d\cos\theta_k d\cos\theta_l d\phi} \cos^2\theta_k \cos\theta_l = \frac{S_{6s}}{10} \tag{7}$$

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}sin^{2}\theta_{k}sin^{2}\theta_{l}cos2\phi = \frac{8S_{3}}{25}$$
 (8)

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}sin2\theta_{k}sin2\theta_{k}cos\phi = \frac{8S_{4}}{25}$$
 (9)

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}sin2\theta_{k}sin\theta_{l}cos\phi = \frac{2S_{5}}{5}$$
 (10)

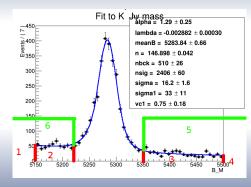
$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}sin2\theta_{k}sin\theta_{l}sin\phi = \frac{2S_{7}}{5}$$
(11)

$$\frac{d^{3}\Gamma}{\Gamma dcos\theta_{k}dcos\theta_{l}d\phi}sin2\theta_{k}sin2\theta_{l}sin\phi = \frac{8S_{8}}{25}$$
 (12)



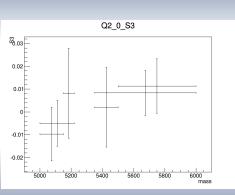
Studied background region

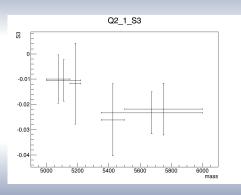
- Defined B⁰ mass bins: $1: (5,5,15) \cup 2: (5.15,5.22) \cup 3: (5.35,5.5) \cup 4: (5.5,6)$ GeV
 - Region 5:(5.35, 6)
 - 2 Region 6:(5, 5.22)
- use the old a^2 bins:
 - 0:0,2
 - 1:2, 4.3
 - 2:4.3, 8, 6
 - 3:10.1, 12.9
 - 4:14.2, 16
 - 5:16, 19



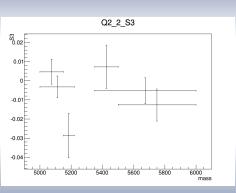
Please remember the numbers, we will need then later on.

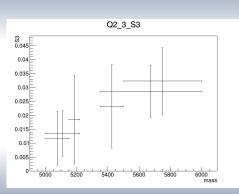




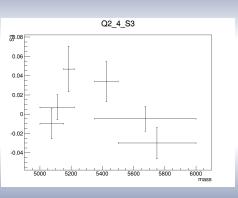


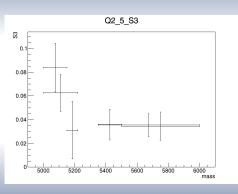




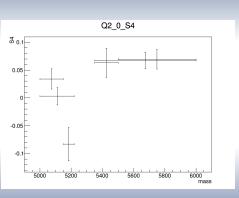


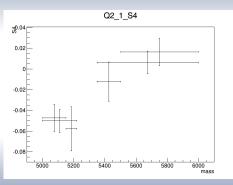




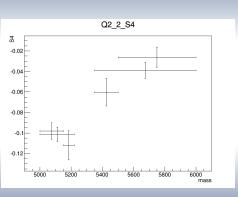


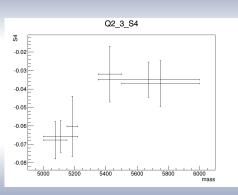




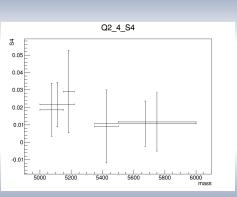


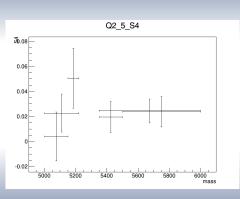




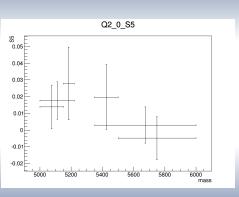


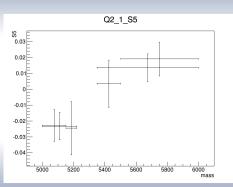




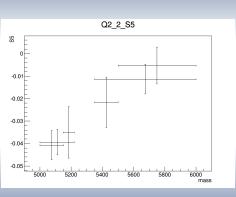


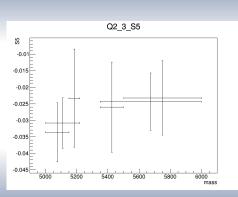




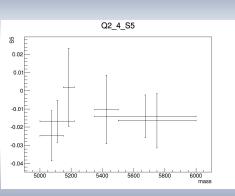


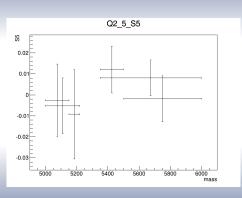




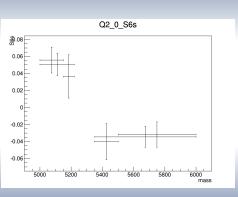


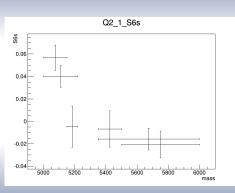




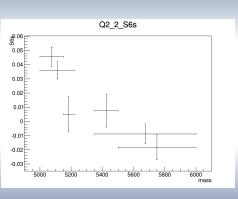


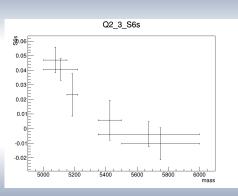




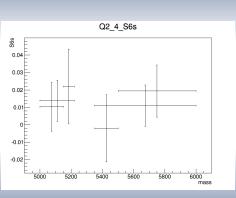


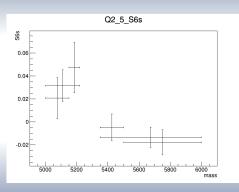




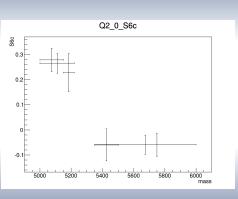


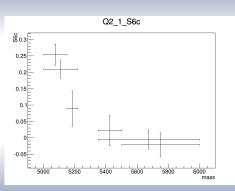




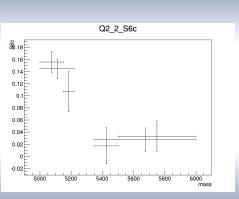


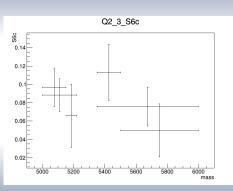




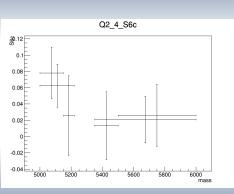


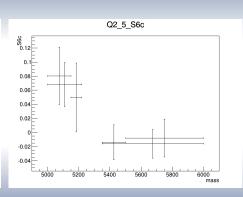




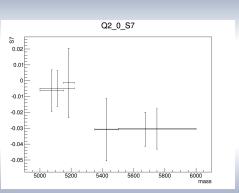


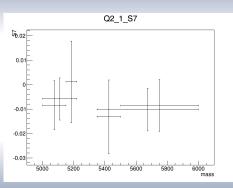




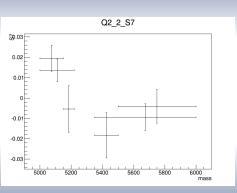


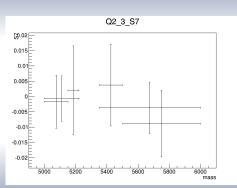




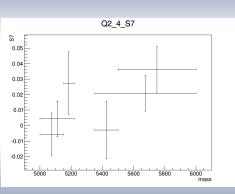


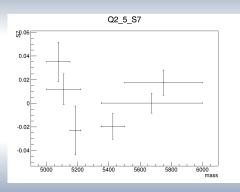




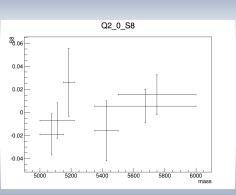


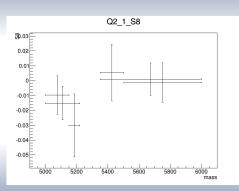




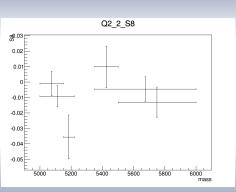


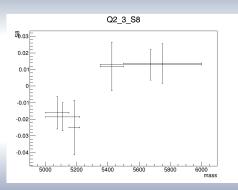




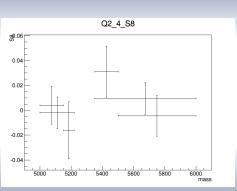


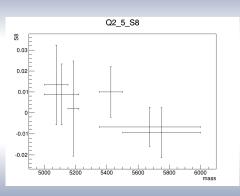




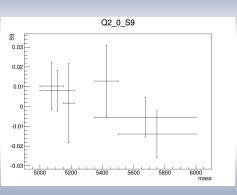


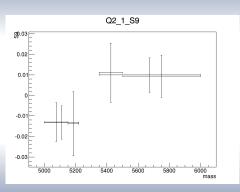




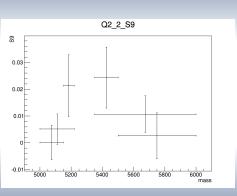


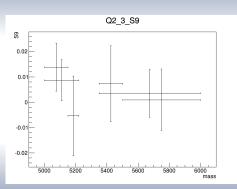




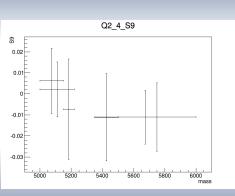


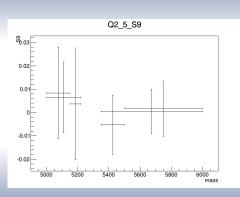




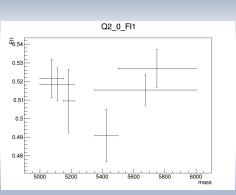


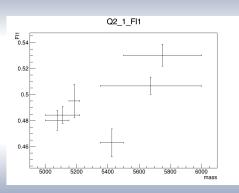




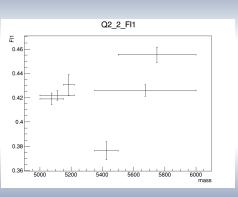


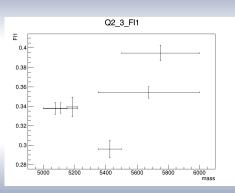




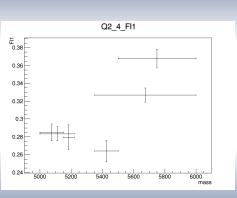


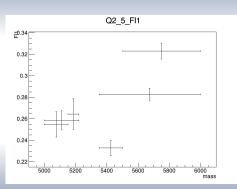




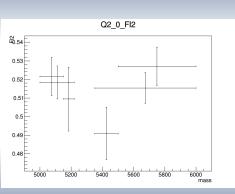


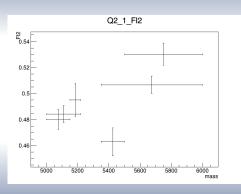




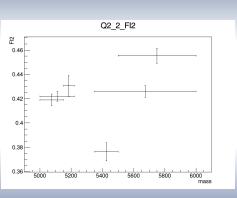


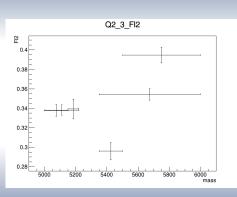




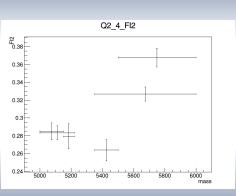


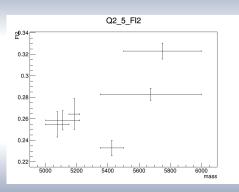




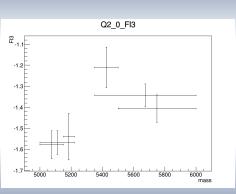


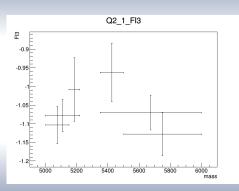




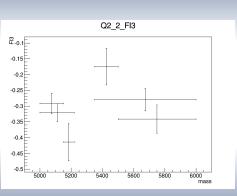


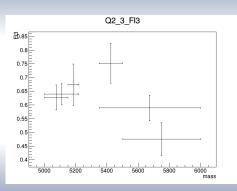




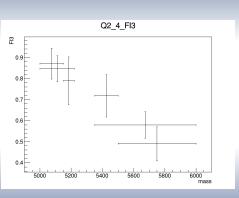


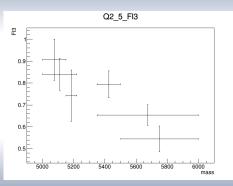




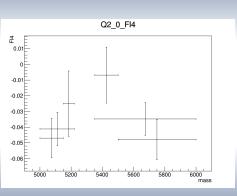


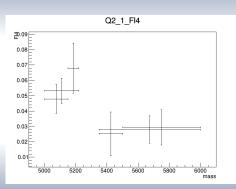




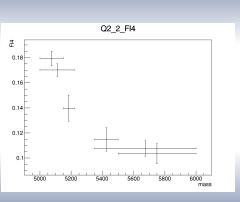


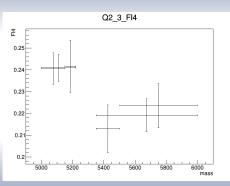




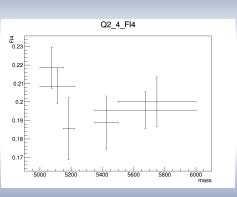


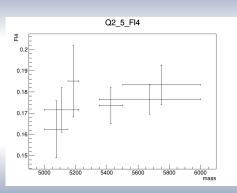














Summary

- Background moments are effectivelly 0
- Apart from F_I and S_6
- S_6 is sizeable at the left hand sideband for certain bins, evidence of partially reconstructed semileptonic decays?
- Because the moments are small, they should have small effect on the final result :)

Wish list:

- Repeat the same with smaller q^2 bins.
- ullet Optimise the binning in q^2 taking into account background systematics and error on signal
- Do unfolding.