Recent results on angular analysis of decay $B o K^* \mu \mu$ at LHCb





University of Padova March 22, 2016

Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

Outline

FIXME!

- 1. Why flavour is important.
- 2. $b \rightarrow s\ell\ell$ theory in a nutshell.
- 3. LHCb measurements of $b \rightarrow s\ell\ell$.
- 4. Global fit to $b \to s \ell \ell$ measurements.
- 5. Conclusions.

A lesson from history - GIM mechanism



- Cabibbo angle was successful in explaining dozens of decay rates in the 1960s.
- There was, however, one that was not observed by experiments: $K^0 \rightarrow \mu^- \mu^+$.
- Glashow, Iliopoulos, Maiani (GIM) mechanism was proposed in the 1970 to fix this problem. The mechanism required the existence of the 4th quark.
- At that point most of the people were skeptical about that. Fortunately in 1974 the discovery of the J/ψ meson silenced the skeptics.



A lesson from history - CKM matrix



- Similarly CP violation was discovered in 1960s in the neutral kaons decays.
- 2×2 Cabbibo matrix could not allow for any CP violation.
- For the CP violation to be possible one needs at least a 3 × 3 unitary matrix → Cabibbo-Kobayashi-Maskawa matrix (1973).
- It predicts existence of *b* (1977) and *t* (1995) quarks.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

A lesson from history - Weak neutral current



- Weak neutral currents were first, introduced in 1958 by Buldman.
- Later on they were naturally incorporated into unification of weak and electromagnetic interactions.
- 't Hooft proved that the GWS models was renormalizable.
- Everything was there on theory side, only missing piece was the experiment, till 1973.



LHCb detector - tracking



- Excellent Impact Parameter (IP) resolution (20 μ m). \Rightarrow Identify secondary vertices from heavy flavour decays
- Proper time resolution $\sim~40~{\rm fs.}$
 - \Rightarrow Good separation of primary and secondary vertices.
- Excellent momentum ($\delta p/p \sim 0.4 0.6\%$) and inv. mass resolution. \Rightarrow Low combinatorial background.

p

 $L \sim 7 \,\mathrm{mm} \mathrm{SV}$

LHCb detector - particle identification





- Excellent Muon identification $\epsilon_{\mu
 ightarrow \mu} \sim 97\%$, $\epsilon_{\pi
 ightarrow \mu} \sim 1-3\%$
- Good $K \pi$ separation via RICH detectors, $\epsilon_{K \to K} \sim 95\%$, $\epsilon_{\pi \to K} \sim 5\%$. \Rightarrow Reject peaking backgrounds.
- High trigger efficiencies, low momentum thresholds. Muons: $p_T > 1.76 \text{GeV}$ at L0, $p_T > 1.0 \text{GeV}$ at HLT1, $B \rightarrow J/\psi X$: Trigger $\sim 90\%$.

Why rare decays?

- In SM allows only the charged interactions to change flavour.
 - Other interactions are flavour conserving.
- One can escape this constrain and produce $b \to s$ and $b \to d$ at loop level.
 - $\circ~$ This kind of processes are suppressed in SM \rightarrow Rare decays.
 - New Physics can enter in the loops.



 W^{\pm}

Tools in rare B^0 decays

Operator Product Expansion and Effective Field Theory

$$H_{eff} = -\frac{4G_f}{\sqrt{2}}VV'^* \sum_{i} \left[\underbrace{\underbrace{C_i(\mu)O_i(\mu)}_{\text{left-handed}} + \underbrace{C_i'(\mu)O_i'(\mu)}_{\text{right-handed}}}_{\text{right-handed}} \right], \qquad \begin{array}{c} \text{i=1.2 Iree} \\ \text{i=3-6.8 Gluon penguin} \\ \text{i=7 Photon penguin} \\ \text{i=5 Scalar penguin} \\ \text{i=5 Scalar penguin} \\ \text{i=P pre-inducted penguin} \\ \text{i=P pre-inducte$$

where C_i are the Wilson coefficients and O_i are the corresponding effective operators.



$B \to K^* \mu \mu$ as a golden channel

Marcin Chrząszcz (Universität Zürich, IFJ PAN)

⁹/39

$B^0 \rightarrow K^* \mu^- \mu^+$ kinematics

 \Rightarrow The kinematics of $B^0 \to K^* \mu^- \mu^+$ decay is described by three angles θ_l , θ_k , ϕ and invariant mass of the dimuon system (q^2) .

 $\Rightarrow \cos \theta_k$: the angle between the direction of the kaon in the K^* $(\overline{K^*})$ rest frame and the direction of the K^* $(\overline{K^*})$ in the B^0 $(\overline{B}{}^0)$ rest frame.

 $\Rightarrow \cos \theta_l$: the angle between the direction of the μ^- (μ^+) in the dimuon rest frame and the direction of the dimuon in the B^0 (\overline{B}^0) rest frame.

⇒ ϕ : the angle between the plane containing the μ^- and μ^+ and the plane containing the kaon and pion from the K^* .



$B^0 \rightarrow K^* \mu^- \mu^+$ kinematics

⇒ The kinematics of $B^0 \to K^* \mu^- \mu^+$ decay is described by three angles θ_l , θ_k , ϕ and invariant mass of the dimuon system (q^2).

$$\frac{d^{4}\Gamma}{dq^{2} \operatorname{dcos} \theta_{K} \operatorname{dcos} \theta_{l} \operatorname{d}\phi} = \frac{9}{32\pi} \left[J_{1s} \sin^{2} \theta_{K} + J_{1c} \cos^{2} \theta_{K} + (J_{2s} \sin^{2} \theta_{K} + J_{2c} \cos^{2} \theta_{K}) \cos 2\theta_{l} \right. \\ \left. + J_{3} \sin^{2} \theta_{K} \sin^{2} \theta_{l} \cos 2\phi + J_{4} \sin 2\theta_{K} \sin 2\theta_{l} \cos \phi + J_{5} \sin 2\theta_{K} \sin \theta_{l} \cos \phi \right. \\ \left. + (J_{6s} \sin^{2} \theta_{K} + J_{6c} \cos^{2} \theta_{K}) \cos \theta_{l} + J_{7} \sin 2\theta_{K} \sin \theta_{l} \sin \phi + J_{8} \sin 2\theta_{K} \sin 2\theta_{l} \sin \phi \right. \\ \left. + J_{9} \sin^{2} \theta_{K} \sin^{2} \theta_{l} \sin 2\phi \right],$$

$$(1)$$

 \Rightarrow This is the most general expression of this kind of decay.

Transversity amplitudes

 \Rightarrow One can link the angular observables to transversity amplitudes

$$\begin{split} J_{1s} &= \frac{(2+\beta_{\ell}^2)}{4} \left[|A_{\perp}^L|^2 + |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right] + \frac{4m_{\ell}^2}{q^2} \operatorname{Re} \left(A_{\perp}^L A_{\perp}^{R*} + A_{\parallel}^L A_{\parallel}^{R*} \right) \,, \\ J_{1c} &= |A_0^L|^2 + |A_0^R|^2 + \frac{4m_{\ell}^2}{q^2} \left[|A_t|^2 + 2\operatorname{Re}(A_0^L A_0^{R*}) \right] + \beta_{\ell}^2 |A_S|^2 \,, \\ J_{2s} &= \frac{\beta_{\ell}^2}{4} \left[|A_{\perp}^L|^2 + |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right] \,, \qquad J_{2c} = -\beta_{\ell}^2 \left[|A_0^L|^2 + |A_0^R|^2 \right] \,, \\ J_3 &= \frac{1}{2} \beta_{\ell}^2 \left[|A_{\perp}^L|^2 - |A_{\parallel}^L|^2 + |A_{\perp}^R|^2 - |A_{\parallel}^R|^2 \right] \,, \qquad J_4 = \frac{1}{\sqrt{2}} \beta_{\ell}^2 \left[\operatorname{Re}(A_0^L A_{\parallel}^{L*} + A_0^R A_{\parallel}^{R*}) \right] \,, \\ J_5 &= \sqrt{2} \beta_{\ell} \left[\operatorname{Re}(A_0^L A_{\perp}^{L*} - A_0^R A_{\perp}^{R*}) - \frac{m_{\ell}}{\sqrt{q^2}} \operatorname{Re}(A_{\parallel}^L A_{S}^* + A_{\parallel}^{R*} A_{S}) \right] \,, \\ J_{6s} &= 2\beta_{\ell} \left[\operatorname{Re}(A_{\parallel}^L A_{\perp}^{L*} - A_{\parallel}^R A_{\perp}^{R*}) \right] \,, \qquad J_{6c} = 4\beta_{\ell} \, \frac{m_{\ell}}{\sqrt{q^2}} \operatorname{Re}(A_0^L A_{S}^* + A_0^{R*} A_{S}) \end{split}$$

$$J_7 \quad = \quad \sqrt{2}\beta_\ell \left[\mathrm{Im}(\mathbf{A}_0^{\mathrm{L}}\mathbf{A}_\parallel^{\mathrm{L}*} - \mathbf{A}_0^{\mathrm{R}}\mathbf{A}_\parallel^{\mathrm{R}*}) + \frac{\mathbf{m}_\ell}{\sqrt{\mathbf{q}^2}} \operatorname{Im}(\mathbf{A}_\perp^{\mathrm{L}}\mathbf{A}_{\mathrm{S}}^* - \mathbf{A}_\perp^{\mathrm{R}*}\mathbf{A}_{\mathrm{S}})) \right],$$

 $J_8 = \frac{1}{\sqrt{2}} \beta_\ell^2 \left[\operatorname{Im}(A_0^L A_\perp^{L^*} + A_0^R A_\perp^{R^*}) \right], \qquad \qquad J_9 = \beta_\ell^2 \left[\operatorname{Im}(A_\parallel^{L^*} A_\perp^{L} + A_\parallel^{R^*} A_\perp^{R}) \right], \qquad \qquad J_9 = \beta_\ell^2 \left[\operatorname{Im}(A_\parallel^{L^*} A_\perp^{L^*} + A_\parallel^{R^*} A_\perp^{R^*}) \right],$

Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Link to effective operators

 \Rightarrow So here is where the magic happens. At leading order the amplitudes can be written as:

$$A_{\perp}^{L,R} = \sqrt{2} N m_B (1-\hat{s}) \left[(\mathcal{C}_9^{\text{eff}} + \mathcal{C}_9^{\text{eff}}) \mp (\mathcal{C}_{10} + \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\text{eff}} + \mathcal{C}_7^{\text{eff}}) \right] \xi_{\perp}(E_{K^*})$$

$$A_{\parallel}^{L,R} = -\sqrt{2}Nm_{B}(1-\hat{s}) \left[(\mathcal{C}_{9}^{\text{eff}} - \mathcal{C}_{9}^{\text{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + \frac{2\hat{m}_{b}}{\hat{s}} (\mathcal{C}_{7}^{\text{eff}} - \mathcal{C}_{7}^{\text{eff}}) \right] \xi_{\perp}(E_{K^{*}})$$

$$A_{0}^{L,R} = -\frac{Nm_{B}(1-\hat{s})^{2}}{2\hat{m}_{K^{*}}\sqrt{\hat{s}}} \left[(\mathcal{C}_{9}^{\text{eff}} - \mathcal{C}_{9}^{\text{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + 2\hat{m}_{b}(\mathcal{C}_{7}^{\text{eff}} - \mathcal{C}_{7}^{\text{eff}}) \right] \xi_{\parallel}(E_{K^{*}}), \quad (3)$$

where $\hat{s} = q^2/m_B^2$, $\hat{m}_i = m_i/m_B$. The $\xi_{\parallel,\perp}$ are the form factors.

Link to effective operators

 \Rightarrow So here is where the magic happens. At leading order the amplitudes can be written as:

$$A_{\perp}^{L,R} = \sqrt{2} N m_B (1-\hat{s}) \left[(\mathcal{C}_9^{\text{eff}} + \mathcal{C}_9^{\text{eff}}) \mp (\mathcal{C}_{10} + \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\text{eff}} + \mathcal{C}_7^{\text{eff}}) \right] \xi_{\perp} (E_{K^*})$$

$$A_{\parallel}^{L,R} = -\sqrt{2}Nm_B(1-\hat{s}) \left[(\mathcal{C}_9^{\rm eff} - \mathcal{C}_9^{\rm eff}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\rm eff} - \mathcal{C}_7^{\rm eff}) \right] \xi_{\perp}(E_{K^*})$$

$$A_{0}^{L,R} = -\frac{Nm_{B}(1-\hat{s})^{2}}{2\hat{m}_{K^{*}}\sqrt{\hat{s}}} \left[(\mathcal{C}_{9}^{\text{eff}} - \mathcal{C}_{9}^{\text{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + 2\hat{m}_{b}(\mathcal{C}_{7}^{\text{eff}} - \mathcal{C}_{7}^{\text{eff}}) \right] \xi_{\parallel}(E_{K^{*}}), \quad (3)$$

where $\hat{s} = q^2/m_B^2$, $\hat{m}_i = m_i/m_B$. The $\xi_{\parallel,\perp}$ are the form factors. \Rightarrow Now we can construct observables that cancel the ξ form factors at leading order:

$$P_5' = \frac{J_5 + \bar{J}_5}{2\sqrt{-(J_2^c + \bar{J}_2^c)(J_2^s + \bar{J}_2^s)}}$$
(4)

²/39

LHCbs $B^0 \rightarrow K^* \mu^- \mu^+$, Selection

 \Rightarrow Trigger

- Muon trigger.
- Topological trigger.
- \Rightarrow Good modelling with MC.
- \Rightarrow Selection:
- As loose as possible.
- Based on the B⁰ vertex quality, impact parameters, loose Particle identification for the hadrons.
- The variables were chosen in a way we are sure the are correctly modelled in MC.



³/39

Peaking backgrounds

 \Rightarrow A number of peaking backgrounds that can mistaken as your signal.

 \Rightarrow There where a specially designed vetoes to fight each of them.

	after preselection, b	efore vetoes	after vetoes and selection		
Channel	Estimated events	% signal	Estimated events	% signal	
$\Lambda_b \rightarrow \Lambda^* (1520)^0 \mu \mu$	$(1.0 \pm 0.5) \times 10^3$	19 ± 8	51 ± 25	1.0 ± 0.4	
$\Lambda_h \rightarrow p K \mu \mu$	$(1.0 \pm 0.5) \times 10^2$	1.9 ± 0.8	5.7 ± 2.8	0.11 ± 0.05	
$B_d^0 \to K^+ \mu \mu$	28 ± 7	0.55 ± 0.06	1.6 ± 0.5	0.031 ± 0.006	
$B_s^0 \to \phi \mu \mu$	$(3.2 \pm 1.3) \times 10^2$	6.2 ± 2.1	17 ± 7	0.33 ± 0.12	
signal swaps	$(3.6 \pm 0.9) \times 10^2$	6.9 ± 0.6	33 ± 9	0.64 ± 0.06	
$B^0_d ightarrow K^* J/\psi$ swaps	$(1.3 \pm 0.4) \times 10^{2}$	2.6 ± 0.4	2.7 ± 2.8	0.05 ± 0.05	



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

¹⁴/39

Multivariate simulation

- PID, kinematics and isolation variables used in a Boosted Decision Tree (BDT) to discriminate signal and background.
- BDT with k-Folding technique.
- Completely data drive.





Multivariate simulation, efficiency

 \Rightarrow BDT was also check in order not to bias our angular distribution:



 \Rightarrow The BDT has small impact on our angular observables. We will correct for this effects later on.

Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

¹⁶/39

Mass modelling

 \Rightarrow The signal is modelled by a sum of two Crystal-Ball function with common mean.

- ⇒ The background is a single exponent
- \Rightarrow The base parameters is performed in the proxy channel:

 $B_d^0 \to J/\psi(\mu\mu)K^*$.

- \Rightarrow All the parameters are fixed in the signal pdf.
- \Rightarrow Scaling factors for resolution are determined from MC.

 \Rightarrow In fitting the rare mode only the signal, background yield and the slope of the exponential is left floating.







⇒ The S-wave fraction is extracted using a LASS model.
Particle Phenomenology, Particle Astrophysics and Cosmology Seminar



Monte Carlo corrections

⇒ No Monte Carlo simulation is perfect! One needs to correct for remaining differences. ⇒ We reweighed our $B_d^0 \to K^* \mu \mu$ Monte Carlo accordingly to differences between the $B_d^0 \to K^* J/\psi$ in data (Splot) and Monte Carlo.



Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

Monte Carlo corrections

⇒ No Monte Carlo simulation is perfect! One needs to correct for remaining differences. ⇒ We reweighed our $B_d^0 \to K^* \mu \mu$ Monte Carlo accordingly to differences between the $B_d^0 \to K^* J/\psi$ in data (Splot) and Monte Carlo.



Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

Detector acceptance

- Detector distorts our angular distribution.
- We need to model this effect.
- 4D function is used:

 $\epsilon(\cos\theta_l,\cos\theta_k,\phi,q^2) = \sum_{ijkl} P_i(\cos\theta_l) P_j(\cos\theta_k) P_k(\phi) P_l(q^2),$

where P_i is the Legendre polynomial of order i.

- We use up to $4^{th}, 5^{th}, 6^{th}, 5^{th}$ order for the $\cos \theta_l, \cos \theta_k, \phi, q^2$.
- The coefficients were determined using Method of Moments, with a huge simulation sample.
- The simulation was done assuming a flat phase space and reweighing the q² distribution to make is flat.
- To make this work the *q*² distribution needs to be reweighed to be flat.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Detector acceptance

- Detector distorts our angular distribution.
- We need to model this effect.
- 4D function is used:

$$\epsilon(\cos\theta_l,\cos\theta_k,\phi,q^2) = \sum_{ijkl} P_i(\cos\theta_l) P_j(\cos\theta_k) P_k(\phi) P_l(q^2),$$

where P_i is the Legendre polynomial of order i.

- We use up to $4^{th}, 5^{th}, 6^{th}, 5^{th}$ order for the $\cos \theta_l, \cos \theta_k, \phi, q^2$.
- The coefficients were determined using Method of Moments, with a huge simulation sample.
- The simulation was done assuming a flat phase space and reweighing the q² distribution to make is flat.
- To make this work the *q*² distribution needs to be reweighed to be flat.





Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Control channel

- We tested our unfolding procedure on $B \to J/\psi K^*.$
- The result is in perfect agreement with other experiments and our different analysis of this decay.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

The columns of New Physics



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

The columns of New Physics

- 1. Maximum likelihood fit:
 - $\circ~$ The most standard way of obtaining the parameters.
 - Suffers from convergence problems, under coverages, etc. in low statistics.
- 2. Method of moments:
 - $\circ~$ Less precise then the likelihood estimator (10-15% larger uncertainties).
 - $\circ~$ Does not suffer from the problems of likelihood fit.
- 3. Amplitude fit:
 - Incorporates all the physical symmetries inside the amplitudes! The most precise estimator.
 - Has theoretical assumptions inside!

Maximum likelihood fit

 $\Rightarrow \text{ In the maximum likelihood fit one could weight the events accordingly}$ $to the <math>\frac{1}{\varepsilon(\cos\theta_l,\cos\theta_k,\phi,q^2)}$

 \Rightarrow Better alternative is to put the efficiency into the maximum likelihood fit itself:

$$\mathcal{L} = \prod_{i=1}^{N} \epsilon_i(\Omega_i, q_i^2) \mathcal{P}(\Omega_i, q_i^2) / \int \epsilon(\Omega, q^2) \mathcal{P}(\Omega, q^2) d\Omega dq^2$$

 \Rightarrow Only the relative weights matters!

 \Rightarrow The Procedure was commissioned with TOY MC study.

 \Rightarrow Angular background component is modelled with 2^{nd} order Chebyshev polynomials, which was tested on the side-bands.

\Rightarrow See arXiv::1503.04100, F.Beaujean , M.Chrzaszcz, N.Serra, D. van Dyk for details.



Results in $B \to K^* \mu \mu$



- Tension with 3 fb^{-1} gets confirmed!
- The two bins deviate both in $2.8~\sigma$ from SM prediction.
- Result compatible with previous result.

Branching fraction measurements of $B \rightarrow K^{*\pm} \mu \mu$



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

°/39

Branching fraction measurements of $B_s^0 \rightarrow \phi \mu \mu$



- Recent LHCb measurement [JHEPP09 (2015) 179].
- Suppressed by $\frac{f_s}{f_d}$.
- Cleaner because of narrow ϕ resonance.
- 3.3σ deviation in SM in the $1-6 {
 m GeV}^2$ bin.

Branching fraction measurements of $\Lambda_b \rightarrow \Lambda \mu \mu$



- This years LHCb measurement [JHEP 06 (2015) 115]].
- In total ~ 300 candidates in data set.
- Decay not present in the low q^2 .

Branching fraction measurements of $\Lambda_b \rightarrow \Lambda \mu \mu$



- This years LHCb measurement [JHEP 06 (2015) 115]].
- In total ~ 300 candidates in data set.
- Decay not present in the low q^2 .

Angular analysis of $\Lambda_b \rightarrow \Lambda \mu \mu$

• For the bins in which we have $> 3 \sigma$ significance the forward backward asymmetry for the hadronic and leptonic system.



- A_{FB}^{H} is in good agreement with SM.
- A_{FB}^{ℓ} always in above SM prediction.

Lepton universality test

- If Z' is responsible for the P'_5 anomaly, does it couple equally to all flavours? $R_{\rm K} = \frac{\int_{q^2=1}^{q^2=6\,{\rm GeV}^2/c^4} ({\rm d}\mathcal{B}[B^+ \to K^+\mu^+\mu^-]/{\rm d}q^2){\rm d}q^2}{\int_{q^2=1}^{q^2=6\,{\rm GeV}^2/c^4} ({\rm d}\mathcal{B}[B^+ \to K^+e^+e^-]/{\rm d}q^2){\rm d}q^2} = 1 \pm \mathcal{O}(10^{-3}) \ .$
- Challenging analysis due to bremsstrahlung.
- Migration of events modeled by MC.
- Correct for bremsstrahlung.
- Take double ratio with $B^+ \rightarrow J/\psi K^+$ to cancel systematics.
- In 3fb⁻¹, LHCb measures $R_K = 0.745^{+0.090}_{-0.074}(stat.)^{+0.036}_{-0.036}(syst.)$
- Consistent with SM at 2.6σ .



 Phys. Rev. Lett. 113, 151601 (2014)



Angular analysis of $B^0 \rightarrow K^* ee$

- With the full data set $(3fb^{-1})$ we performed angular analysis in $0.0004 < q^2 < 1 \ {\rm GeV}^2$.
- Electrons channels are extremely challenging experimentally:
 - Bremsstrahlung.
 - Trigger efficiencies.
- Determine the angular observables: $F_{
 m L}$, $A_{
 m T}^{
 m (2)}$, $A_{
 m T}^{
 m Re}$, $A_{
 m T}^{
 m Im}$:

$$\begin{split} F_{\rm L} &= \frac{|A_0|^2}{|A_0|^2 + |A_{||}|^2 + |A_{\perp}|^2} \\ A_{\rm T}^{(2)} &= \frac{|A_{\perp}|^2 - |A_{||}|^2}{|A_{\perp}|^2 + |A_{||}|^2} \\ A_{\rm T}^{\rm Re} &= \frac{2\mathcal{R}e(A_{||L}A_{\perp L}^* + A_{||R}A_{\perp R}^*)}{|A_{||}|^2 + |A_{\perp}|^2} \\ A_{\rm T}^{\rm Im} &= \frac{2\mathcal{I}m(A_{||L}A_{\perp L}^* + A_{||R}A_{\perp R}^*)}{|A_{||}|^2 + |A_{\perp}|^2}, \end{split}$$

⁸¹/39

(5

Angular analysis of $B^0 \rightarrow K^* ee$



- Results in full agreement with the SM.
- Similar strength on C_7 Wilson coefficient as from $b \rightarrow s\gamma$ decays.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

³¹/₃₉



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

- A preliminary fit prepared by S. Descotes-Genon, L. Hofer, J. Matias, J. Virto, presented at 1510.04239
- Took into the fit:
 - $\circ~\mathcal{B}(B \rightarrow X_s \gamma) = (3.36 \pm 0.23) \times 10^{-4}$, Misiak et. al. 2015.
 - $\circ~\mathcal{B}(B\to\mu\mu)$, theory: Bobeth et al 2013, experiment: LHCb+CMS average (2015)
 - $\circ \ \mathcal{B}(B
 ightarrow X_s \mu \mu)$, Huber et al 2015
 - $\circ \ \mathcal{B}(B
 ightarrow K \mu \mu)$,Bouchard et al 2013, 2015
 - $\circ \ PB_{(s)} \rightarrow K^*(\phi) \mu \mu$, Horgan et al 2013
 - $\circ B \rightarrow Kee$, $B \rightarrow K^*ee$ and R_k .
- Overall there is around 4.5σ discrepancy wrt. SM.



- A preliminary fit prepared by S. Descotes-Genon, L. Hofer, J. Matias, J. Virto, presented at 1510.04239
- The data can be explained by modifying the C_9 Wilson coefficient.
- Overall there is around $4.5 \; \sigma$ discrepancy wrt. SM.



³/39

Coefficient	Best fit	1σ	3σ	$\mathrm{Pull}_{\mathrm{SM}}$	p-value (%)
$\mathcal{C}_7^{\mathrm{NP}}$	-0.02	[-0.04, -0.00]	[-0.07, 0.04]	1.1	16.0
$\mathcal{C}_9^{ m NP}$	-1.11	[-1.32, -0.89]	[-1.71, -0.40]	4.5	62.0
$\mathcal{C}_{10}^{\mathrm{NP}}$	0.58	[0.34, 0.84]	[-0.11, 1.41]	2.5	25.0
$\mathcal{C}^{\mathrm{NP}}_{7'}$	0.02	[-0.01, 0.04]	[-0.05, 0.09]	0.7	15.0
$\mathcal{C}_{9'}^{\mathrm{NP}}$	0.49	[0.21, 0.77]	[-0.33, 1.35]	1.8	19.0
$\mathcal{C}^{\mathrm{NP}}_{10'}$	-0.27	[-0.46, -0.08]	[-0.84, 0.28]	1.4	17.0
$\mathcal{C}_9^{\rm NP}=\mathcal{C}_{10}^{\rm NP}$	-0.21	[-0.40, 0.00]	[-0.74, 0.55]	1.0	16.0
$\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{10}^{\rm NP}$	-0.69	[-0.88, -0.51]	[-1.27, -0.18]	4.1	55.0
$\mathcal{C}_{9'}^{\rm NP}=\mathcal{C}_{10'}^{\rm NP}$	-0.09	[-0.35, 0.17]	[-0.88, 0.66]	0.3	14.0
$\mathcal{C}_{9'}^{\rm NP} = -\mathcal{C}_{10'}^{\rm NP}$	0.20	[0.08, 0.32]	[-0.15, 0.56]	1.7	19.0
$\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{9'}^{\rm NP}$	-1.09	[-1.28, -0.88]	[-1.62, -0.42]	4.8	72.0
$\begin{aligned} \mathcal{C}_9^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= -\mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned}$	-0.68	[-0.49, -0.49]	[-1.36, -0.15]	3.9	50.0
$ \begin{aligned} \mathcal{C}_9^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= \mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned} $	-0.17	[-0.29, -0.06]	[-0.54, 0.18]	1.5	18.0

Table 2: Best-fit points, confidence intervals, pulls for the SM hypothesis and p-values for different one-dimensional NP scenarios.

Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

If not NP?

- We are not there yet!
- There might be something not taken into account in the theory.
- Resonances ($J\!/\!\psi$, $\psi(2S)$) tails can mimic NP effects.
- There might be some non factorizable QCD corrections. "However, the central value of this effect would have to be significantly larger than expected on the basis of existing estimates" D.Straub, 1503.06199.



If not NP?

- We are not there yet!
- There might be something not taken into account in the theory.
- Resonances (J/ψ , $\psi(2S)$) tails can mimic NP effects.
- There might be some non factorizable QCD corrections. "However, the central value of this effect would have to be significantly larger than expected on the basis of existing estimates" D.Straub, 1503.06199.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

If not NP?

- How about our clean P_i observables?
- The QCD cancel as mentioned only at leading order.
- Comparison to normal observables with the optimised ones.



There is more!

• There is one other LUV decay recently measured by LHCb.

•
$$R(D^*) = \frac{\mathcal{B}(B \to D^* \tau \nu)}{\mathcal{B}(B \to D^* \mu \nu)}$$

- Clean SM prediction: $R(D^*) = 0.252(3)$, PRD 85 094025 (2012)
- • LHCb result: $R(D^*)=0.336\pm 0.027\pm 0.030,$ HFAG average: $R(D^*)=0.322\pm 0.022$
- 3.9σ discrepancy wrt. SM.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

Conclusions

- Clear tensions wrt. SM predictions!
- Measurements cluster in the same direction.
- We are not opening the champagne yet!
- Still need improvement both on theory and experimental side.
- Time will tell if this is QCD+fluctuations or new Physics:

Conclusions

- Clear tensions wrt. SM predictions!
- Measurements cluster in the same direction.
- We are not opening the champagne yet!
- Still need improvement both on theory and experimental side.
- Time will tell if this is QCD+fluctuations or new Physics:

"... when you have eliminated all the Standard Model explanations, whatever remains, however improbable, must be New Physics." prof. Joaquim Matias

Thank you for the attention!



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

³⁹/39

Backup



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

Particle Phenomenology, Particle Astrophysics and Cosmology Seminar

A lesson from history - GIM mechanism



- Cabibbo angle was successful in explaining dozens of decay rates in the 1960s.
- There was, however, one that was not observed by experiments: $K^0 \rightarrow \mu^- \mu^+$.
- Glashow, Iliopoulos, Maiani (GIM) mechanism was proposed in the 1970 to fix this problem. The mechanism required the existence of the 4th quark.
- At that point most of the people were skeptical about that. Fortunately in 1974 the discovery of the J/ψ meson silenced the skeptics.



A lesson from history - CKM matrix



- Similarly CP violation was discovered in 1960s in the neutral kaons decays.
- 2×2 Cabbibo matrix could not allow for any CP violation.
- For the CP violation to be possible one needs at least a 3 × 3 unitary matrix
 ↔ Cabibbo-Kobayashi-Maskawa matrix (1973).
- It predicts existence of *b* (1977) and *t* (1995) quarks.



Marcin Chrząszcz (Universität Zürich, IFJ PAN)

A lesson from history - Weak neutral current



- Weak neutral currents were first, introduced in 1958 by Buldman.
- Later on they were naturally incorporated into unification of weak and electromagnetic interactions.
- 't Hooft proved that the GWS models was renormalizable.
- Everything was there on theory side, only missing piece was the experiment, till 1973.



Recent measurements

 \Rightarrow Branching fractions: $B^{0,\pm} \rightarrow K^{0,\pm} \mu^- \mu^+$ LHCb, Mar 14 $B^0 \rightarrow K^* \mu^- \mu^+$ CMS, Jul 15 $B^0_{s} \rightarrow \phi \mu^- \mu^+$ LHCb, Jun 15 $B^{\pm} \rightarrow \pi^{\pm} \mu^{-} \mu^{+}$ LHCb, Sep 15 $\Lambda_b \rightarrow \Lambda \mu^- \mu^+$ LHCb, Mar 15 $B \rightarrow \mu^{-}\mu^{+}$ CMS+LHCb, Jun 15 \Rightarrow CP asymmetry: $B^{\pm} \rightarrow \pi^{\pm} \mu^{-} \mu^{+}$ LHCb, Sep 15 \Rightarrow lsospin asymmetry: $B \rightarrow K \mu^{-} \mu^{+}$ LHCb, Mar 14

 $\begin{array}{l} \Rightarrow \mbox{Lepton Universality:} \\ B^{\pm} \rightarrow K^{\pm} \ell \bar{\ell} & \mbox{LHCb, Jun 14} \\ \Rightarrow \mbox{Angular:} \\ B^{0} \rightarrow K^{*} \ell \bar{\ell} & \mbox{LHCb, Jan 15} \\ B^{\pm} \rightarrow K^{*,\pm} \ell \bar{\ell} & \mbox{BaBar, Aug 15} \\ B^{0}_{s} \rightarrow \phi \ell \bar{\ell} & \mbox{LHCb, Jun 15} \\ \Lambda_{b} \rightarrow \Lambda \mu^{-} \mu^{+} & \mbox{LHCb, Mar 15} \end{array}$

Recent measurements

 \Rightarrow Branching fractions: $B^{0,\pm}
ightarrow K^{0,\pm} \mu^- \mu^+$ LHCb, Mar 14 $B^0 \rightarrow K^* \mu^- \mu^+$ CMS, Jul 15 $B_{\rm s}^0 \rightarrow \phi \mu^- \mu^+$ LHCb, Jun 15 $B^{\pm} \rightarrow \pi^{\pm} \mu^{-} \mu^{+}$ LHCb, Sep 15 $\Lambda_b \rightarrow \Lambda \mu^- \mu^+$ LHCb, Mar 15 $B \rightarrow \mu^{-}\mu^{+}$ CMS+LHCb, Jun 15 \Rightarrow CP asymmetry: $B^{\pm} \rightarrow \pi^{\pm} \mu^{-} \mu^{+}$ LHCb, Sep 15 \Rightarrow lsospin asymmetry: $B \rightarrow K \mu^{-} \mu^{+}$ LHCb, Mar 14

 $\begin{array}{l} \Longrightarrow \text{ Lepton Universality:} \\ B^{\pm} \to K^{\pm} \ell \overline{\ell} & \text{LHCb, Jun 14} \\ \Longrightarrow \text{ Angular:} \\ B^{0} \to K^{*} \ell \overline{\ell} & \text{LHCb, Jan 15} \\ B^{\pm} \to K^{*,\pm} \ell \overline{\ell} & \text{BaBar, Aug 15} \\ B^{0}_{s} \to \phi \ell \overline{\ell} & \text{LHCb, Jun 15} \\ \Lambda_{b} \to \Lambda \mu^{-} \mu^{+} & \text{LHCb, Mar 15} \end{array}$

$>2~\sigma$ deviations from SM