# *інср*

# Overview of recent experimental results in flavour physics

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# Flavour Physics, WHAT, WHY HOW?

- ⇒ WHAT: Quarks and leptons exists in 6 "flavors" (u,c,t,d,s,b) and (e, $\mu$ ,  $\tau$ ,  $\nu_e$ ,  $\nu_{\mu}$ ,  $\nu_{\tau}$ ). ⇒ WHY:
- Flavour is the heart of SM. It involves 22 from 28 free parameters, like masses mixing and CP violation.
- Flavor physics loop processes (box and penguins) are sensitive to energy scales well beyond the ones of the accelerators, thanks to virtual contributions.



#### $\rightarrow$ Indirect search for New Physics

- $\Rightarrow$  HOW:
- Compare precise theoretical predictions with precise experimental measurements.
- LHCb, Belle, BaBar, ATLAS, CMS, NA62, BESIII, neutrinos experiments,...

### Introduction to flavor physics

 $\Rightarrow$  Masses and mixing of quarks have a common origin in the SM: The Yukawa interactions with the Higgs:

$$\mathcal{L}_{\text{Yukawa}} = -Y_{ij}^d \bar{Q'}_{Li} \phi d'_{Rj} - Y_{ij}^u \bar{Q'}_{Ri} \phi u'_{Rj}$$

 $\Rightarrow$  The masses are generated using SSB by diagonalizing Y. The CKM matrix relats the mass eigenstates with the flavour eigenstates:

$$\begin{pmatrix} d \\ s \\ b \end{pmatrix}' = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}^{\text{phys}}$$

 $\Rightarrow$  The charge current interactions between quarks are proportional to the CKM matrix elements:

$$\mathcal{L}_{CC} = -\frac{g}{\sqrt{2}}\bar{u}\gamma^{\mu}(1-\gamma^5)V_{CKM}d + \dots$$



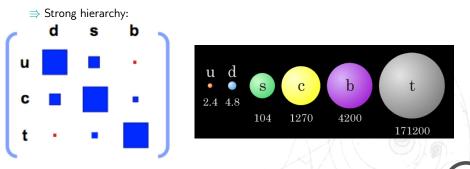
## **CP** violation

 $\Rightarrow$  The  $3\times3$  CKM has build inside the CP violation:

$$V_{ij} \neq V_{ij}^* \Rightarrow (CP)\mathcal{L}_{CC} \neq \mathcal{L}^{\dagger}$$

 $\Rightarrow$  In the Wolfenstein parametrization the CKM matrix reads:

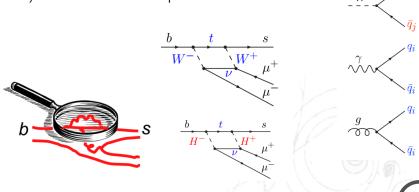
$$V_{CKM} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix}$$



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# Why rare decays?

- In SM allows only the charged interactions to change flavour.
  - Other interactions are flavour conserving.
- One can escape this constrain and produce  $b \to s$  and  $b \to d$  at loop level.
  - $\circ~$  This kind of processes are suppressed in SM  $\rightarrow$  Rare decays.
  - New Physics can enter in the loops.



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 $W^{\pm}$ 

# Searching for New Physics

- $\Rightarrow$  The fundamental questions:
- Why 3 generations? Why such hierarchy structure?
- Stability of the Higgs vacum? Dark Matter?
- Baryon asymmetry of the universe? CP in SM is to small!

# Searching for New Physics

- $\Rightarrow$  The fundamental questions:
- Why 3 generations? Why such hierarchy structure?
- Stability of the Higgs vacum? Dark Matter?
- Baryon asymmetry of the universe? CP in SM is to small!
- $\Rightarrow$  Two ways to answer them:
- Direct searches: try to produce directly new real particles "on-shell", but we don't know their mass or lifetime and we are limited by the center-of-mass energy of accelerator.
- Indirect searches: study the effect of "off-shell" (virtual) particles within quantum loop. Compare precise theoretical predictions with precise experimental measurements. Not limited by the center-of-mass energy of accelerator. It happened in the past:
  - $\circ$  CP violation in the Kaon system: existence of b and t quarks.
  - $\circ~$  Lack of observation of  $K^0_S \to \mu \mu$ : existence of c quark.
  - $\circ~$  Neutral weak currents: existence of Z boson.
- Very powerful tool!

# Selected physics results:

- Rare Decays
  - $\begin{array}{l} \circ \ B^0_s/B^0_d \to \mu\mu \\ \circ \ B^0_d \to K^*\mu\mu, \ B^0_s \to \phi\mu\mu, \ \Lambda_b \to \Lambda\mu\mu. \end{array}$
- Tests of lepton universalities:

$$\circ R_k = \mathcal{B}(B^+ \to K^+ \mu \mu) / \mathcal{B}(B^+ \to K^+ ee)$$
  
 
$$\circ R(D), R(D^*)$$

- $(g-2)_{\mu}$
- CP violation:
  - $\circ \ \gamma$  angle.
  - $\circ$  CP violation in  $B_d^0$  and  $B_s^0$ .
  - CP violation in charm.
  - $\circ~$  CP violation in kaons.
  - $\circ \ V_{ub}.$

# Heavy flavor experiments

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# Rare decays

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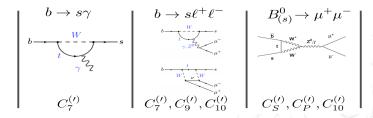
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### Tools

#### • Operator Product Expansion and Effective Field Theory

$$H_{eff} = -\frac{4G_f}{\sqrt{2}}VV'^* \sum_{i} \left[\underbrace{\underbrace{C_i(\mu)O_i(\mu)}_{\text{left-handed}} + \underbrace{C_i'(\mu)O_i'(\mu)}_{\text{right-handed}}\right], \qquad \begin{array}{c} \stackrel{\text{i=1,2}}{\underset{i=3-6,8}{\text{Gluon penguin}}} \\ \stackrel{\text{i=3-6,8}}{\underset{i=9.10}{\text{FW penguin}}} \\ \stackrel{\text{i=9,10}}{\underset{i=9}{\text{W penguin}}} \\ \stackrel{\text{i=9,10}}{\underset{i=9}{\text{W$$

where  $C_i$  are the Wilson coefficients and  $O_i$  are the corresponding effective operators.



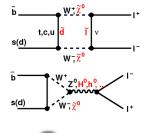
 $B^0 o \mu^+ \mu^-$ 

 Clean theoretical prediction, GIM and helicity suppressed in the SM:

• 
$$\mathcal{B}(B^0_s \to \mu^- \mu^+) = (3.65 \pm 0.23) \times 10^{-9}$$

• 
$$\mathcal{B}(B^0 \to \mu^- \mu^+) = (1.06 \pm 0.09) \times 10^{-10}$$

- Sensitive to contributions from scalar and pesudoscalar couplings.
- Probing: MSSM, higgs sector, etc.
- In MSSM:  ${\cal B}(B^0_s\to\mu^-\mu^+)\sim {\rm tg}^6\,\beta/m_A^4$



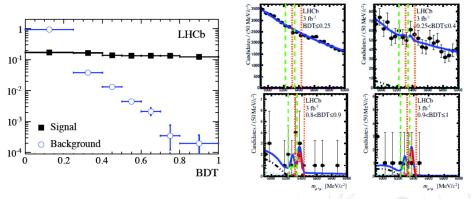


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# $B^0 \rightarrow \mu^+ \mu^-$ searches

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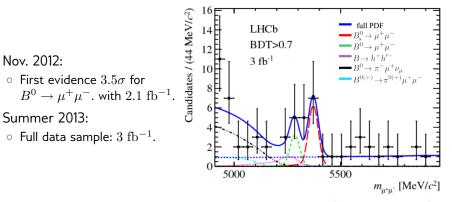
 Background rejection power is a key feature of rare decays → use multivariate classifiers (BDT) and strong PID.



• Normalize the BF to  $B^+ \to J/\psi(\mu\mu)K^+$  and  $B^0 \to K\pi$ .

# $B^0 \rightarrow \mu^+ \mu^-$ Results

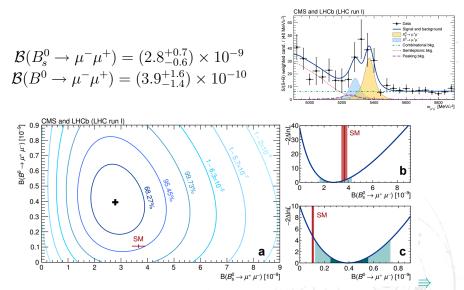
Nov. 2012:



- Measured BF:  $\mathcal{B}(B^0_s \to \mu^- \mu^+) = (2.9^{+1.1}_{-1.0}(stat.)^{+0.3}_{-0.1}(syst.)) \times 10^{-9}$
- 4.0σ significance!
- $\mathcal{B}(B^0 \to \mu^- \mu^+) < 7 \times 10^{-10}$  at 95% CL
- CMS result: PRL 111 (2013) 101805

# LHCb+CMS Combination

#### Nature 522 (2015) 68



#### $2.3 \; \sigma$ compatibility with SM!

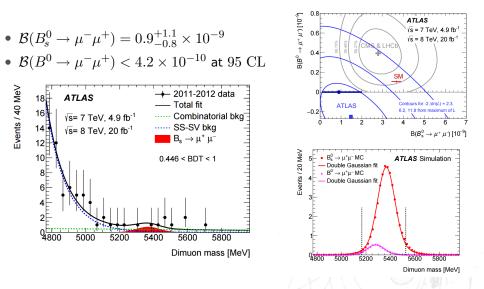
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### ATLAS enters the game

#### arXiv:1604.04263

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# $B_d^0 \to K^* \mu \mu$

⇒ The decay of  $B_d^0 \to K^* \mu \mu$  has number of angular observables that are sensitive to different Wilson coefficients:  $C_7^{(\prime)}$ ,  $C_9^{(\prime)}$ ,  $C_{10}^{(\prime)}$ . ⇒ The complete angular expression is given by:

$$\frac{1}{\mathrm{d}\Gamma/\mathrm{d}q^2} \frac{\mathrm{d}^4\Gamma}{\mathrm{d}\cos\theta_\ell\,\mathrm{d}\cos\theta_K\,\mathrm{d}\phi\,\mathrm{d}q^2} = \frac{9}{32\pi} \left[ \frac{3}{4} (1-F_\mathrm{L}) \sin^2\theta_K + F_\mathrm{L}\cos^2\theta_K + \frac{1}{4} (1-F_\mathrm{L}) \sin^2\theta_K\cos2\theta_\ell \right]$$
$$- F_\mathrm{L}\cos^2\theta_K\cos2\theta_\ell + S_3\sin^2\theta_K\sin^2\theta_\ell\cos2\phi + S_4\sin2\theta_\ell\cos\varphi + S_4\sin2\theta_\ell\cos\varphi + S_5\sin2\theta_K\sin^2\theta_\ell\cos\varphi + S_6\sin^2\theta_K\cos\varphi + S_6\sin^2\theta_K\cos\varphi + S_6\sin^2\theta_K\sin^2\theta_\ell\sin\varphi + S_6\sin^2\theta_K\sin^2\theta_\ell\sin\varphi + S_8\sin^2\theta_K\sin^2\theta_\ell\sin\varphi \right]$$

 $\Rightarrow$  Furthermore, one can construct a form factor free observables:

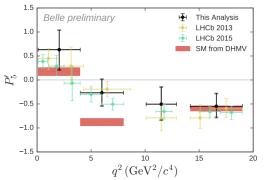
$$P_5' = \frac{S_5}{F_L(1 - F_L)}$$

- $\Rightarrow$  Analysis performed with 3 methods:
- Likelihood fit.
- Method of moments.
- Amplitudes fit.

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# $B^0_d \rightarrow K^* \mu \mu$ results

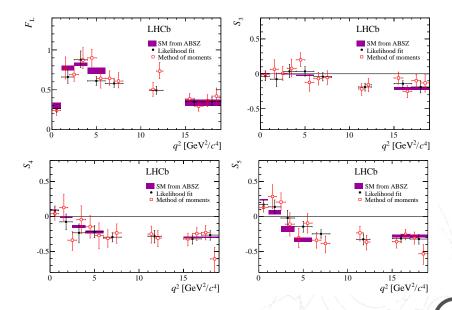
arxiv:1604.04042



- Tension with 3 fb<sup>-1</sup> gets confirmed!
- two bins both deviate by  $2.8 \sigma$  from SM prediction.
- Result compatible with previous results and Belle!
- SM: JHEP12(2014)125

# $\overline{B^0_d \to K^* \mu \mu}$ results

#### JHEP 02 (2016) 104

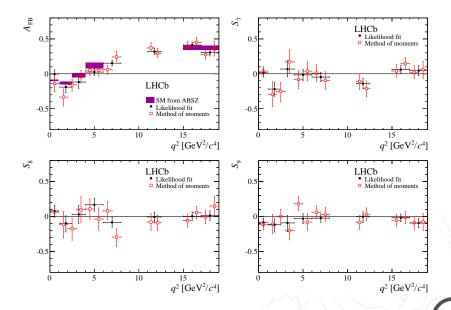


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# $\overline{B^0_d \to K^* \mu \mu}$ results

#### JHEP 02 (2016) 104



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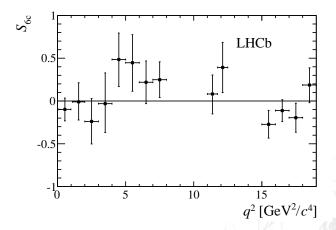
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# $B_d^0 \to K^* \mu \mu$ results

JHEP 02 (2016) 104

 $\Rightarrow$  Method of Moments allowed us to measure for the first time a new observable:



 $\Rightarrow$  LHCb also measured the CP asymmetries with Method of Moments and the likelihood fit that are consistent with SM

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## Compatibility with SM

#### JHEP 02 (2016) 104

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⇒ Use EOS software package to test compatibility with SM. ⇒ Perform the  $\chi^2$  fit to the measured:

$$F_L, A_{FB}, S_{3,...,9}.$$

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 $\Rightarrow \text{Float a vector coupling:} \\ \Re(C_9).$ 

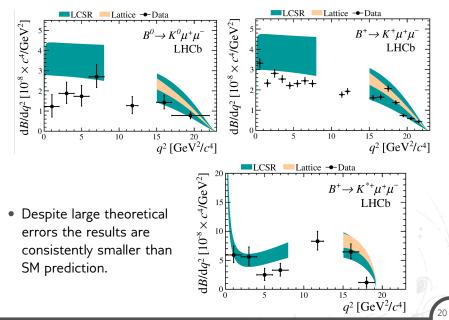
 $\Rightarrow$  Best fit is found to be 3.4  $\sigma$  away from the SM.

$$\Delta \Re(C_9) \equiv \Re(C_9)^{\text{III}} - \Re(C_9)^{\text{SM}} = -1.03$$

C.

# BF of $B \to K^{*\pm} \mu \mu$

#### JHEP 07 (2012) 133

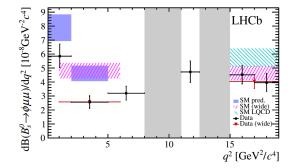


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#### JHEP09 (2015) 179

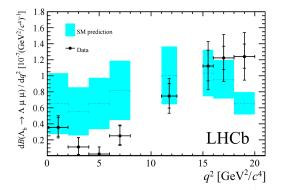


- Last years LHCb measurement.
- Suppressed by  $\frac{f_s}{f_d}$ .
- Cleaner because of narrow  $\phi$  resonance.
- $3.3 \sigma$  deviation in SM in the  $1 6 GeV^2$  bin.

#### JHEP 06 (2015) 115

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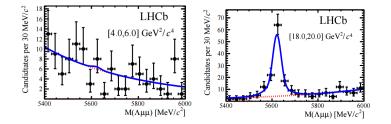
# BF of $\Lambda_{\!b} \to \Lambda \mu \mu$



- Last years LHCb measurement.
- In total  $\sim 300$  candidates in data set.
- Decay not present in the low  $q^2$ .

# BF of $\Lambda_b \to \Lambda \mu \mu$

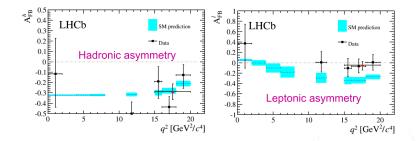
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- Last years LHCb measurement.
- In total  $\sim 300$  candidates in data set.
- Decay not present in the low  $q^2$ .

# Angular analysis of $\Lambda_b \rightarrow \Lambda \mu \mu$

• For the bins in which we have  $> 3 \sigma$  significance the forward backward asymmetry for the hadronic and leptonic system.



- $A_{FB}^{H}$  is in good agreement with SM.
- $A_{FB}^{\ell}$  always in above SM prediction.

JHEP 06 (2015) 115

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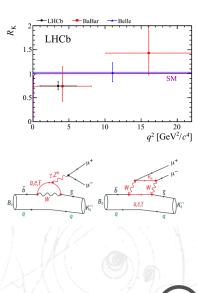
# Lepton Universality tests

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### Lepton universality test

- Does the NP couple equally to all flavours?  $R_{\rm K} = \frac{\int_{q^2=1}^{q^2=6\,{\rm GeV}^2/c^4} ({\rm d}\mathcal{B}[B^+ \to K^+\mu^+\mu^-]/{\rm d}q^2){\rm d}q^2}{\int_{q^2=1}^{q^2=6\,{\rm GeV}^2/c^4} ({\rm d}\mathcal{B}[B^+ \to K^+e^+e^-]/{\rm d}q^2){\rm d}q^2} = 1 \pm \mathcal{O}(10^{-3}) .$
- Challenging analysis due to bremsstrahlung.
- Migration of events modeled by MC.
- Correct for bremsstrahlung.
- Take double ratio with  $B^+ \rightarrow J/\psi K^+$  to cancel systematics.
- In  $3 \text{fb}^{-1}$ , LHCb measures:  $R_K = 0.745^{+0.090}_{-0.074}(stat.)^{+0.036}_{-0.036}(syst.)$
- Consistent with SM at  $2.6\sigma$ .

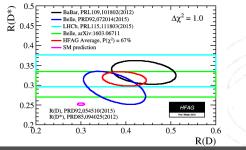


### More Lepton universality tests

• There is one other LUV decay recently measured by LHCb.

• 
$$R(D^*) = \frac{\mathcal{B}(B \to D^* \tau \nu)}{\mathcal{B}(B \to D^* \mu \nu)}$$

- Clean SM prediction:  $R(D^*) = 0.252(3)$ , PRD 85 094025 (2012)
- • LHCb result:  $R(D^*)=0.336\pm 0.027\pm 0.030,$  HFAG average:  $R(D^*)=0.322\pm 0.022$
- $4.0 \sigma$  discrepancy wrt. SM.



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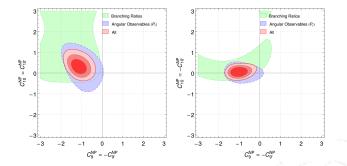
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## Explanation of anomalies

#### arXiv:1510.04239

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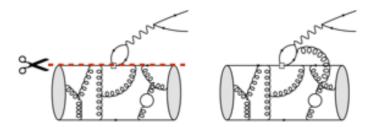
 $\Rightarrow$  Thanks to S. Descotes-Genon, L.Hofer, J.Matias, J.Virto we have a global fit to the anomalies.



⇒ The fit prefer a modification of  $C_9$  Wilson coefficient with a value of  $C_9^{NP} = -1$ , with a significance over  $4\sigma$ . ⇒ Many theories link might accommodate the observed deviations.

### Explanation of anomalies

- We are not there yet!
- There might be something not taken into account in the theory.
- Resonances ( $J\!/\!\psi$ ,  $\psi(2S)$ ) tails can mimic NP effects.
- There might be some non factorizable QCD corrections. "However, the central value of this effect would have to be significantly larger than expected on the basis of existing estimates" D.Straub, 1503.06199.



 $(g-2)\mu$ 

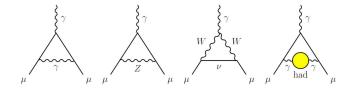
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# Muon anomalous magnetic moment $(g-2)_{\mu}$

 $\Rightarrow$  Dirac equations predict a muon magnetic moment  $\vec{M} = g_{\mu} \frac{e}{2m_{\mu}} \vec{S}$ with a gyromagnetic ratio  $g_{\mu} = 2$ . Experimentally  $g_{\mu} > 2$ .

 $\Rightarrow$  This anomaly  $a_{\mu}$  arises from calculable quantum fluctuations:

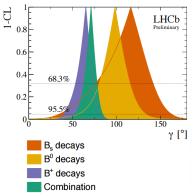


 $\Rightarrow \mathsf{SM} \text{ value } a_{\mu}^{\mathrm{SM}} = (116591803 \pm 49) \times 10^{-11} \text{ [PDG 2014]}$   $\Rightarrow \mathsf{Experiment: } a_{\mu}^{\mathrm{E821}} = (116592091 \pm 63) \times 10^{-11} \text{ [PDG 2014]}$   $\Rightarrow \mathsf{Difference: } a_{\mu}^{\mathrm{E821}} - a_{\mu}^{\mathrm{SM}} = (288 \pm 80) \times 10^{-11} \Leftrightarrow 3.6 \ \sigma!$   $\Rightarrow \mathsf{Underestimated uncertainties? \mathsf{SUSY? NP?} }$   $\Rightarrow \mathsf{Fermilabs E989 should futher reduce the experimental error to } \pm 16 \times 10^{-11}.$ 

# $\gamma$ angle

$$\Rightarrow \gamma = \arg\left(-\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*}\right) \approx \arg\left(-\frac{V_{ub}^*}{V_{cb}^*}\right) \text{ is the last well known CKM angles.}$$

- $\Rightarrow$  Can be determined by:
- Tree level processes, nearly insensitive to NP. Act as reference. Negligible theoretical uncertainty, using B → DK.
- Loop processes, sensitive to NP.
- ⇒ Comparing the two can reveal NP!



- $\Rightarrow$  The LHCb combinations leads to:  $\gamma = 70.9^{+7.1}_{-8.5}$ .
- $\Rightarrow$  Improved by  $2^{\rm deg}$  compared to our previous result!

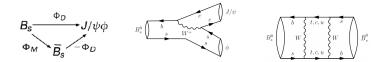
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- ⇒ Other B-factories have 2 times larger errors:
- BaBar:  $\gamma = (70 \pm 18)^{\text{deg}}$
- Belle:  $\gamma = (73^{+18}_{-15})^{\text{deg}}$

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# Mixing induced CPV in $B_s^0$

 $\Rightarrow$  Interference between  $B^0_s$  decaying to  $J/\psi\phi$  either directly or by oscillations gives rise to CP violation phase:  $\phi_s^{J/\psi\phi}$ 



 $\begin{array}{l} \Rightarrow \text{ In the Sm } \phi_s \approx -2\beta + s = -(0.0376^{+0.0007}_{-0.0008}) \text{ rad, where } \beta_s = \\ \arg\left(-\frac{V_{ts}V^*_{tb}}{V_{cs}V^*_{cb}}\right) \\ \Rightarrow \text{ At the leading order the same phase is expected } B^0_s \rightarrow D_s D_s \text{ and} \end{array}$ 

- $B \to J/\psi \pi \pi$ .
- $\Rightarrow$  NP can enter in the  $B_s^0$  mixing!
- $\Rightarrow$  Measured by simultaneous fit to  $B_s^0$  and  $\bar{B_s^0}$  decay rates:

$$\frac{\mathrm{d}^{4}\Gamma(B_{s}^{0}\to J/\psi\phi)}{\mathrm{d}t\,\mathrm{d}\cos\theta_{\mu}\,\mathrm{d}\varphi_{h}\,\mathrm{d}\cos\theta_{K}}=f(\phi_{s},\Delta\Gamma_{s},\Gamma_{s},\Delta m_{s},\mathcal{M}(B_{s}^{0}),|\mathcal{A}_{\perp}|,|\mathcal{A}_{\parallel}|,|\mathcal{A}_{S}|,\delta_{\perp},\delta_{\parallel},\ldots)$$

# Backup

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# Theory implications

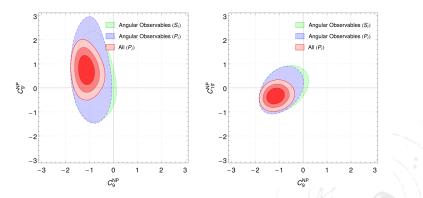
| Coefficient   | Best fit | $1\sigma$      | $3\sigma$      | $\mathrm{Pull}_{\mathrm{SM}}$ | p-value (%) |
|---|----------|----------------|----------------|-------------------------------|-------------|
| $\mathcal{C}_7^{\mathrm{NP}}$   | -0.02    | [-0.04, -0.00] | [-0.07, 0.04]  | 1.1                           | 16.0        |
| $\mathcal{C}_9^{\mathrm{NP}}$   | -1.11    | [-1.32, -0.89] | [-1.71, -0.40] | 4.5                           | 62.0        |
| $\mathcal{C}_{10}^{\mathrm{NP}}$  | 0.58     | [0.34, 0.84]   | [-0.11, 1.41]  | 2.5                           | 25.0        |
| $\mathcal{C}^{\mathrm{NP}}_{7'}$  | 0.02     | [-0.01, 0.04]  | [-0.05, 0.09]  | 0.7                           | 15.0        |
| $\mathcal{C}^{\mathrm{NP}}_{9'}$  | 0.49     | [0.21, 0.77]   | [-0.33, 1.35]  | 1.8                           | 19.0        |
| $\mathcal{C}^{\mathrm{NP}}_{10'}$   | -0.27    | [-0.46, -0.08] | [-0.84, 0.28]  | 1.4                           | 17.0        |
| $\mathcal{C}_9^{\rm NP}=\mathcal{C}_{10}^{\rm NP}$  | -0.21    | [-0.40, 0.00]  | [-0.74, 0.55]  | 1.0                           | 16.0        |
| $\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{10}^{\rm NP}$   | -0.69    | [-0.88, -0.51] | [-1.27, -0.18] | 4.1                           | 55.0        |
| $\mathcal{C}_{9'}^{\rm NP}=\mathcal{C}_{10'}^{\rm NP}$  | -0.09    | [-0.35, 0.17]  | [-0.88, 0.66]  | 0.3                           | 14.0        |
| $\mathcal{C}_{9'}^{\rm NP} = -\mathcal{C}_{10'}^{\rm NP}$   | 0.20     | [0.08, 0.32]   | [-0.15, 0.56]  | 1.7                           | 19.0        |
| $\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{9'}^{\rm NP}$   | -1.09    | [-1.28, -0.88] | [-1.62, -0.42] | 4.8                           | 72.0        |
| $ \begin{aligned} \mathcal{C}_9^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= -\mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned} $ | -0.68    | [-0.49, -0.49] | [-1.36, -0.15] | 3.9                           | 50.0        |
| $\begin{aligned} \mathcal{C}_{9}^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= \mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned}$  | -0.17    | [-0.29, -0.06] | [-0.54, 0.18]  | 1.5                           | 18.0        |

Table 2: Best-fit points, confidence intervals, pulls for the SM hypothesis and p-values for different one-dimensional NP scenarios.

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### If not NP?

- How about our clean  $P_i$  observables?
- The QCD cancel as mentioned only at leading order.
- Comparison to normal observables with the optimised ones.



### Transversity amplitudes

 $\Rightarrow$  One can link the angular observables to transversity amplitudes

$$\begin{split} J_{1s} &= \frac{(2+\beta_{\ell}^2)}{4} \left[ |A_{\perp}^L|^2 + |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right] + \frac{4m_{\ell}^2}{q^2} \operatorname{Re} \left( A_{\perp}^L A_{\perp}^{R*} + A_{\parallel}^L A_{\parallel}^{R*} \right) \,, \\ J_{1c} &= |A_0^L|^2 + |A_0^R|^2 + \frac{4m_{\ell}^2}{q^2} \left[ |A_t|^2 + 2\operatorname{Re}(A_0^L A_0^{R*}) \right] + \beta_{\ell}^2 |A_S|^2 \,, \\ J_{2s} &= \frac{\beta_{\ell}^2}{4} \left[ |A_{\perp}^L|^2 + |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right] \,, \qquad J_{2c} = -\beta_{\ell}^2 \left[ |A_0^L|^2 + |A_0^R|^2 \right] \,, \\ J_3 &= \frac{1}{2} \beta_{\ell}^2 \left[ |A_{\perp}^L|^2 - |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 - |A_{\parallel}^R|^2 \right] \,, \qquad J_4 = \frac{1}{\sqrt{2}} \beta_{\ell}^2 \left[ \operatorname{Re}(A_0^L A_{\parallel}^{L*} + A_0^R A_{\parallel}^{R*}) \right] \,, \\ J_5 &= \sqrt{2} \beta_{\ell} \left[ \operatorname{Re}(A_0^L A_{\perp}^{L*} - A_0^R A_{\perp}^{R*}) - \frac{m_{\ell}}{\sqrt{q^2}} \operatorname{Re}(A_{\parallel}^L A_{S}^* + A_{\parallel}^{R*} A_{S}) \right] \,, \\ J_{6s} &= 2\beta_{\ell} \left[ \operatorname{Re}(A_{\parallel}^L A_{\perp}^{L*} - A_{\parallel}^R A_{\perp}^{R*}) \right] \,, \qquad J_{6c} = 4\beta_{\ell} \, \frac{m_{\ell}}{\sqrt{q^2}} \operatorname{Re}(A_0^L A_{S}^* + A_0^{R*} A_{S}) \,. \end{split}$$

$$J_7 \quad = \quad \sqrt{2}\beta_\ell \left[ \mathrm{Im}(\mathbf{A}_0^{\mathrm{L}}\mathbf{A}_{\parallel}^{\mathrm{L}*} - \mathbf{A}_0^{\mathrm{R}}\mathbf{A}_{\parallel}^{\mathrm{R}*}) + \frac{\mathbf{m}_\ell}{\sqrt{\mathsf{q}^2}} \, \mathrm{Im}(\mathbf{A}_{\perp}^{\mathrm{L}}\mathbf{A}_{\mathrm{S}}^* - \mathbf{A}_{\perp}^{\mathrm{R}*}\mathbf{A}_{\mathrm{S}})) \right],$$

 $J_8 = \frac{1}{\sqrt{2}} \beta_\ell^2 \left[ \operatorname{Im}(\mathbf{A}_0^{\mathbf{L}} \mathbf{A}_\perp^{\mathbf{L}} + \mathbf{A}_0^{\mathbf{R}} \mathbf{A}_\perp^{\mathbf{R}}) \right] , \qquad \qquad J_9 = \beta_\ell^2 \left[ \operatorname{Im}(\mathbf{A}_\parallel^{\mathbf{L}*} \mathbf{A}_\perp^{\mathbf{L}} + \mathbf{A}_\parallel^{\mathbf{R}*} \mathbf{A}_\perp^{\mathbf{R}}) \right] ,$ 

### Link to effective operators

 $\Rightarrow$  So here is where the magic happens. At leading order the amplitudes can be written as (soft form factors):

$$A_{\perp}^{L,R} = \sqrt{2}Nm_B(1-\hat{s}) \left[ (\mathcal{C}_9^{\rm eff} + \mathcal{C}_9^{\rm eff'}) \mp (\mathcal{C}_{10} + \mathcal{C}'_{10}) + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\rm eff} + \mathcal{C}_7^{\rm eff'}) \right] \xi_{\perp}(E_{K^*})$$

$$A_{\parallel}^{L,R} \quad = \quad -\sqrt{2}Nm_B(1-\hat{s})\left[ (\mathcal{C}_9^{\mathrm{eff}} - \mathcal{C}_9^{\mathrm{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\mathrm{eff}} - \mathcal{C}_7^{\mathrm{eff}}) \right] \xi_{\perp}(E_{K^*})$$

$$A_0^{L,R} = -\frac{Nm_B(1-\hat{s})^2}{2\hat{m}_{K^*}\sqrt{\hat{s}}} \left[ (\mathcal{C}_9^{\rm eff} - \mathcal{C}_9^{\rm eff\prime}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + 2\hat{m}_b (\mathcal{C}_7^{\rm eff} - \mathcal{C}_7^{\rm eff\prime}) \right] \xi_{\parallel}(E_{K^*}),$$

where  $\hat{s}=q^2/m_B^2$ ,  $\hat{m}_i=m_i/m_B$ . The  $\xi_{\parallel,\perp}$  are the form factors.

#### Link to effective operators

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$$A_0^{L,R} = -\frac{Nm_B(1-\hat{s})^2}{2\hat{m}_{K^*}\sqrt{\hat{s}}} \left[ (\mathcal{C}_9^{\rm eff} - \mathcal{C}_9^{\rm eff\prime}) \mp (\mathcal{C}_{10} - \mathcal{C}'_{10}) + 2\hat{m}_b (\mathcal{C}_7^{\rm eff} - \mathcal{C}_7^{\rm eff\prime}) \right] \xi_{\parallel}(E_{K^*}),$$

where  $\hat{s} = q^2/m_B^2$ ,  $\hat{m}_i = m_i/m_B$ . The  $\xi_{\parallel,\perp}$  are the form factors.  $\Rightarrow$  Now we can construct observables that cancel the  $\xi$  form factors at leading order:

$$P_5' = \frac{J_5 + J_5}{2\sqrt{-(J_2^c + \bar{J}_2^c)(J_2^s + \bar{J}_2^s)}}$$

# Mass modelling

⇒ The signal is modelled by a sum of two Crystal-Ball functions with common mean. ⇒ The background is a single exponential. ⇒ The base parameters are obtained from the proxy channel:  $B_d^0 \rightarrow J/\psi(\mu\mu)K^*$ . ⇒ All the parameters are fixed in the signal pdf.

 $\Rightarrow$  Scaling factors for resolution are determined from MC.

 $\Rightarrow$  In fitting the rare mode only the signal, background yield and the slope of the exponential is left floating.

 $\Rightarrow$  We found  $624 \pm 30$  candidates in the

most interesting  $[1.1, 6.0]~{\rm GeV^2/c^4}$  region  $\Rightarrow$  The S-wave fraction is extracted using a and  $2398\pm57$  in the full range  $$\rm LASS\ model.$$  [1.1, 19.]  ${\rm GeV^2/c^4}.$$ 

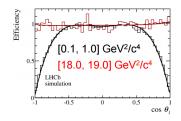
#### Detector acceptance

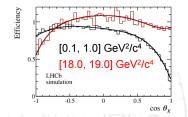
- Detector distorts our angular distribution.
- We need to model this effect.
- 4D function is used:

 $\epsilon(\cos\theta_l,\cos\theta_k,\phi,q^2) = \sum_{ijkl} P_i(\cos\theta_l) P_j(\cos\theta_k) P_k(\phi) P_l(q^2),$ 

where  $P_i$  is the Legendre polynomial of order i.

- We use up to  $4^{th}, 5^{th}, 6^{th}, 5^{th}$  order for the  $\cos \theta_l, \cos \theta_k, \phi, q^2$ .
- The coefficients were determined using Method of Moments, with a huge simulation sample.
- The simulation was done assuming a flat phase space and reweighing the q<sup>2</sup> distribution to make is flat.

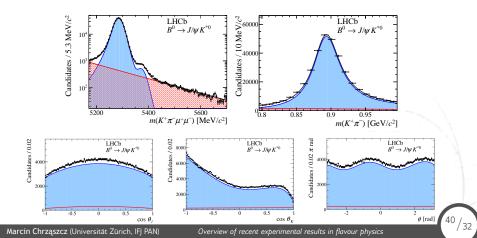




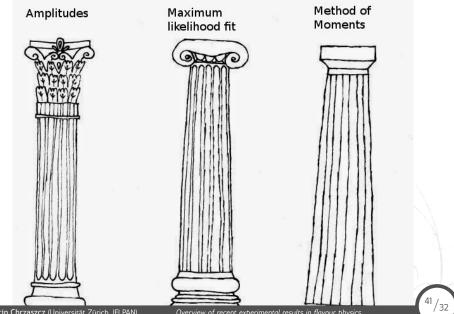
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### Control channel

- We tested our unfolding procedure on  $B \rightarrow J/\psi K^*$ .
- The result is in perfect agreement with other experiments and our different analysis of this decay.



# The columns of New Physics



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Overview of recent experimental results in flavour physics

# $B_d^0 \to K^* \mu \mu$ results

 $\Rightarrow$  In the maximum likelihood fit one could weight the events accordingly to the 1

 $\overline{\varepsilon(\cos\theta_l,\cos\theta_k,\phi,q^2)}$ 

 $\Rightarrow$  Better alternative is to put the efficiency into the maximum likelihood fit itself:

$$\mathcal{L} = \prod_{i=1}^{N} \epsilon_i(\Omega_i, q_i^2) \mathcal{P}(\Omega_i, q_i^2) / \int \epsilon(\Omega, q^2) \mathcal{P}(\Omega, q^2) d\Omega dq^2$$

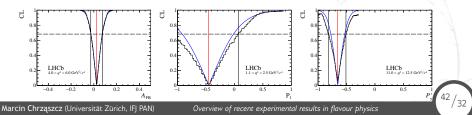
⇒ Only the relative weights matters!

 $\Rightarrow$  The Procedure was commissioned with TOY MC study.

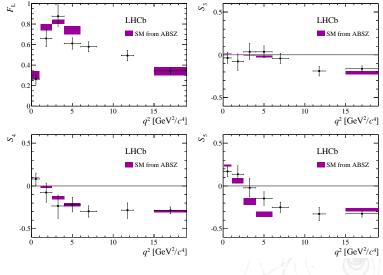
 $\Rightarrow$  Use Feldmann-Cousins to determine the uncertainties.

 $\Rightarrow$  Angular background component is modelled with  $2^{nd}$  order Chebyshev polynomials, which was tested on the side-bands.

 $\Rightarrow$  S-wave component treated as nuisance parameter.



### Maximum likelihood fit - Results

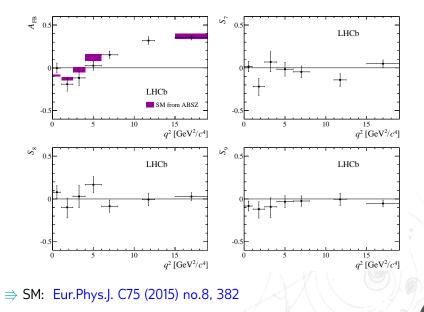


⇒ SM: Eur.Phys.J. C75 (2015) no.8, 382

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### Maximum likelihood fit - Results



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### Method of moments

 $\Rightarrow$  See Phys.Rev.D91(2015)114012, F.Beaujean , M.Chrzaszcz, N.Serra, D. van Dyk for details.

 $\Rightarrow$  The idea behind Method of Moments is simple: Use orthogonality of spherical harmonics,  $f_j(\overrightarrow{\Omega})$  to solve for coefficients within a  $q^2$  bin:

$$\int f_i(\overrightarrow{\Omega}) f_j(\overrightarrow{\Omega}) = \delta_{ij}$$

$$M_i = \int \left(\frac{1}{d(\Gamma + \bar{\Gamma})/dq^2}\right) \frac{d^3(\Gamma + \bar{\Gamma})}{d\overrightarrow{\Omega}} f_i(\overrightarrow{\Omega}) d\Omega$$

 $\Rightarrow$  Don't have true angular distribution but we "sample" it with our data.  $\Rightarrow$  Therefore:  $\int \rightarrow \sum$  and  $M_i \rightarrow \widehat{M}_i$ 

$$\hat{M}_i = \frac{1}{\sum_e \omega_e} \sum_e \omega_e f_i(\overrightarrow{\Omega}_e)$$

 $\Rightarrow$  The weight  $\omega$  accounts for the efficiency. Again the normalization of weights does not matter.

### Amplitudes method

⇒ Fit for amplitudes as (continuous) functions of  $q^2$  in the region:  $q^2 \in [1.1.6.0] \text{ GeV}^2/\text{c}^4$ . ⇒ Needs some Ansatz:

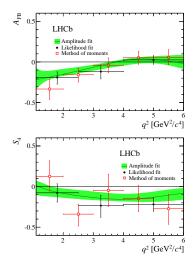
$$A(q^2) = \alpha + \beta q^2 + \frac{\gamma}{q^2}$$

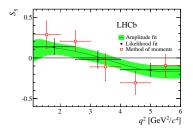
 $\Rightarrow$  The assumption is tested extensively with toys.

- $\Rightarrow$  Set of 3 complex parameters  $\alpha, \beta, \gamma$  per vector amplitude:
- L, R, 0,  $\parallel$ ,  $\perp$ ,  $\Re$ ,  $\Im \rightarrow 3 \times 2 \times 3 \times 2 = 36$  DoF.
- Scalar amplitudes: +4 DoF.
- Symmetries of the amplitudes reduces the total budget to: 28.
- ⇒ The technique is described in JHEP06(2015)084, U. Egede, M. Patel, K.A. Petridis.
- $\Rightarrow$  Allows to build the observables as continuous functions of  $q^2$ :
- At current point the method is limited by statistics.
- In the future the power of this method will increase.

 $\Rightarrow$  Allows to measure the zero-crossing points for free and with smaller errors than previous methods.

### Amplitudes - results





#### Zero crossing points:

| $q_0(S_4) < 2.65$              | at 95% $CL$     |
|--------------------------------|-----------------|
| $q_0(S_5) \in [2.49, 3.95]$    | at $68\% \; CL$ |
| $q_0(A_{FB}) \in [3.40, 4.87]$ | at 68% $CL$     |

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