

# Rare decays at LHCb

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## Rare Decays at LHCb

#### Muonic B decays

 $\begin{array}{l} \Rightarrow & \text{Br } B_s^0/B_d^0 \to \mu\mu/\tau\tau. \\ \Rightarrow & \text{Br + Ang. } B \to K^*\mu\mu. \\ \Rightarrow & \text{Br + Ang. } B_s^0 \to \phi\mu\mu. \\ \Rightarrow & \text{Isospin } B \to K\mu\mu. \\ \Rightarrow & \text{CP asymmetry } B \to \pi\mu\mu. \end{array}$ 

#### Charm decays

 $\Rightarrow D \to \pi \pi \mu \mu$  $\Rightarrow D \to K \pi \mu \mu$  $\Rightarrow D \to e \mu.$ 

⇒ Enormous Physics program
 which is constantly expanding.
 ⇒ Will cover only part of the results.

LFU test  

$$\Rightarrow B^{+} \rightarrow K^{+}\ell\ell$$

$$\Rightarrow B^{0}_{d} \rightarrow K^{*}\ell\ell$$

 $\Rightarrow$  See G.Andreassi talk for LUV!!!

Strange decays

 $\Rightarrow K_{\rm S}^0 \to \mu\mu.$ 

Radiative decays

 $\Rightarrow$  See H.Evans talk.

#### arXiv:1703.05747

# $B_{s/d} \to \mu \mu$

⇒ Golden channel for LHCb. ⇒ Normalized to the  $B \to K\pi$  and  $B \to KJ/\psi$ .

 $\Rightarrow$  The selection is achived by BDT trained on MC and calibrated on data.

$$\Rightarrow \mathcal{B}(B_{s}^{0} \to \mu\mu) = (3.0 \pm 0.6^{+0.3}_{-0.2})10^{-9}$$
  
7.8  $\sigma$  significant!

$$\Rightarrow \mathcal{B}(B^0_d o \mu \mu) < 3.4 imes 10^{-10}$$
, 90%CL

#### Effective lifetime

⇒ Sensitivity to non-scalar NP. ⇒  $\tau(B_s^0 \to \mu\mu) = 2.04 \pm 0.44 \pm 0.05 \text{ps}$ 



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# $B_{s/d} \to \tau \tau$

 $\Rightarrow$  NP sensitivity enhanced due to the high  $\tau$  mass.

 $\Rightarrow$  More challenging: at least  $2\nu$  are escaping.

- $\Rightarrow$  Selecting  $au o 3\pi 
  u$ , o 9.31 %
- $\Rightarrow$  Normalization channel:
- $B \rightarrow D(K\pi\pi)D_{s}(KK\pi).$
- $\Rightarrow$  No peak in the *B* mass window  $\rightarrow$  fit the NN output.





J. High Energy Phys. 04 (2017) 029

#### $\Lambda_b \to p \pi \mu \mu$



- $\Rightarrow$  BDT selection trained on MC
- $\Rightarrow$  Normalized to  $\Lambda_b \rightarrow p\pi J\!/\psi$
- ⇒ With futher QCD improvements we will be able to to measure  $\frac{|V_{ts}|}{|V_{ts}|}$ .





# $\mathcal{B}(\Lambda_b \to p\pi\mu\mu) = (6.9 \pm 1.9 \pm 1.1^{+1.3}_{-1.0}) \times 10^{-8}$

Rare Decays at LHCb

## Search for light scalars

⇒ Hidden sector models are gathering more and more attention.

 $\Rightarrow$  Inflaton model: new scalar then mixes with the Higgs.

 $\Rightarrow$  *B* decays are sensitive as the inflaton might be light.

⇒ Searched for long living particle  $\chi$  produced in:  $B \rightarrow \chi(\mu\mu)K$ .

 $\Rightarrow$  Analysis performed blindly as a peak search.

 $\Rightarrow$  Light inflaton essentially ruled out:





#### Phys. Rev. D 95, 071101 (2017)

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 $K_{S}^{0} \rightarrow \mu \mu$ 

 $\Rightarrow$  *pp* collisions create enormous amount of strange mesons.

 $\Rightarrow$  Can be used to search for  $K_{\rm S}^0 \rightarrow \mu\mu$ .

 $\Rightarrow$  SM prediction:

 $\mathcal{B}(K_{\rm S}^0 \to \mu\mu) = (5.0 \pm 1.5) \times 10^{-12}$ 

 $\Rightarrow$  Dominated by the long distance effects.

 $\Rightarrow$  Bkg dominated by  $K_{S}^{0} \rightarrow \pi\pi$ .





⇒ No significant enhanced of signal has been observed and UL was set:

 $\begin{array}{l} \mathcal{B}(\textit{K}^{\rm 0}_{\rm S} \rightarrow \mu \mu) < 0.8(1.0) \times 10^{-9} \\ {\rm at} \; 90(95)\% \; {\rm CL} \end{array}$ 

#### JHEP 02 (2016) 104, CMS-PAS-BPH-15-008, ATLAS-CONF-2017-023, Phys. Rev. Lett. 118 (2017)

 $\Rightarrow B^0 \rightarrow K^* \mu^- \mu^+$  is a smoking gun for NP hunting!

 $B^0 \rightarrow K^* \mu^- \mu^+$  decay

⇒ Reach angular observables makes
 is sensitive to different NP models
 ⇒ In addition one can construct less
 form factor dependent observables:

$$P_5' = \frac{S_5}{\sqrt{F_L(1 - F_L)}}$$

 $\Rightarrow$  In single analysis observed  $3.4~\sigma$  discrepancy in the  $C_9$  WC.



## Branching fraction measurements of $B_s^0 \rightarrow \phi \mu \mu$



- Recent LHCb measurement, JHEP09 (2015) 179.
- Suppressed by  $\frac{f_s}{f_d}$ .
- Cleaner because of narrow  $\phi$  resonance.
- $3.3 \sigma$  deviation in SM in the  $1 6 \mathrm{GeV}^2$  bin.
- Angular part in agreement with SM ( $S_5$  is not accessible).

## Theory implications of $b \rightarrow s\ell\ell$ JHEP 06 (2016) 092

- A fit prepared by S. Descotes-Genon, L. Hofer, J. Matias, J. Virto.
- The data can be explained by modifying the  $C_9$  Wilson coefficient.
- Overall there is  $> 4 \sigma$  discrepancy wrt. the SM prediction.



## If not NP?

- We are not there yet!
- There might be something not taken into account in the theory.
- Resonances (J/ $\psi$ ,  $\psi(2S)$ ) tails can mimic NP effects.
- There might be some non factorizable QCD corrections. "However, the central value of this effect would have to be significantly larger than expected on the basis of existing estimates" D.Straub, arXiv:1503.06199.



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#### Measurement of phase difference

 $\Rightarrow$  One could try to measure the phase difference between the resonances and the nonresonant amplitudes to see if the interference is large enough to explain the effects.

 $\Rightarrow$  Measured firstly done for the decay  $B \rightarrow K \mu \mu$ .

 $\Rightarrow$  The analysis based:

$$C_9^{\text{eff}} = C_9 + Y(q^2) = C_9 + \sum_j \eta_j e^{i\delta_i} A_j^{\text{res}}(q^2)$$

⇒ The amplitudes are modelled
 Briet-Wigner and Flatte functions.
 ⇒ Interference cannot explain the observed anomalies.



#### Conclusions

- Clear tensions wrt. SM predictions!
- Measurements cluster in the same direction.
- We are not opening the champagne yet!
- Still need improvement both on theory and experimental side.
- Time will tell if this is QCD+fluctuations or new Physics:

#### Conclusions

- Clear tensions wrt. SM predictions!
- Measurements cluster in the same direction.
- We are not opening the champagne yet!
- Still need improvement both on theory and experimental side.
- Time will tell if this is QCD+fluctuations or new Physics:

"... when you have eliminated all the Standard Model explanations, whatever remains, however improbable, must be New Physics." Prof. Joaquim Matias

## Thank you for the attention!



# Backup

<sup>15</sup>/<sub>14</sub>

## Theory implications

Coefficient	Best fit	$1\sigma$	$3\sigma$	$\mathrm{Pull}_{\mathrm{SM}}$	p-value (%)
$\mathcal{C}_7^{\mathrm{NP}}$	-0.02	[-0.04, -0.00]	[-0.07, 0.04]	1.1	16.0
$\mathcal{C}_9^{ m NP}$	-1.11	[-1.32, -0.89]	[-1.71, -0.40]	4.5	62.0
$\mathcal{C}_{10}^{\mathrm{NP}}$	0.58	[0.34, 0.84]	[-0.11, 1.41]	2.5	25.0
$\mathcal{C}^{\mathrm{NP}}_{7'}$	0.02	[-0.01, 0.04]	[-0.05, 0.09]	0.7	15.0
$\mathcal{C}^{\mathrm{NP}}_{9'}$	0.49	[0.21, 0.77]	[-0.33, 1.35]	1.8	19.0
$\mathcal{C}^{\mathrm{NP}}_{10'}$	-0.27	[-0.46, -0.08]	[-0.84, 0.28]	1.4	17.0
$\mathcal{C}_9^{\rm NP}=\mathcal{C}_{10}^{\rm NP}$	-0.21	[-0.40, 0.00]	[-0.74, 0.55]	1.0	16.0
$\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{10}^{\rm NP}$	-0.69	[-0.88, -0.51]	[-1.27, -0.18]	4.1	55.0
$\mathcal{C}_{9'}^{\rm NP}=\mathcal{C}_{10'}^{\rm NP}$	-0.09	[-0.35, 0.17]	[-0.88, 0.66]	0.3	14.0
$\mathcal{C}_{9'}^{\rm NP} = -\mathcal{C}_{10'}^{\rm NP}$	0.20	[0.08, 0.32]	[-0.15, 0.56]	1.7	19.0
$\mathcal{C}_9^{\rm NP} = -\mathcal{C}_{9'}^{\rm NP}$	-1.09	[-1.28, -0.88]	[-1.62, -0.42]	4.8	72.0
$\begin{aligned} \mathcal{C}_9^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= -\mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned}$	-0.68	[-0.49, -0.49]	[-1.36, -0.15]	3.9	50.0
$ \begin{aligned} \mathcal{C}_9^{\mathrm{NP}} &= -\mathcal{C}_{10}^{\mathrm{NP}} \\ &= \mathcal{C}_{9'}^{\mathrm{NP}} = -\mathcal{C}_{10'}^{\mathrm{NP}} \end{aligned} $	-0.17	[-0.29, -0.06]	[-0.54, 0.18]	1.5	18.0

Table 2: Best-fit points, confidence intervals, pulls for the SM hypothesis and p-values for different one-dimensional NP scenarios.

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## If not NP?

- How about our clean  $P_i$  observables?
- The QCD cancel as mentioned only at leading order.
- Comparison to normal observables with the optimised ones.



#### Transversity amplitudes

 $\Rightarrow$  One can link the angular observables to transversity amplitudes

$$J_{1s} \quad = \quad \frac{(2+\beta_{\ell}^2)}{4} \left[ |A_{\perp}^L|^2 + |A_{\parallel}^L|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right] + \frac{4m_{\ell}^2}{q^2} \mathrm{Re} \left( A_{\perp}^L A_{\perp}^{R*} + A_{\parallel}^L A_{\parallel}^{R*} \right) \,,$$

$$J_{1c} \quad = \quad \left|A_0^L\right|^2 + \left|A_0^R\right|^2 + \frac{4m_\ell^2}{q^2} \left[\left|A_t\right|^2 + 2\text{Re}(A_0^L A_0^{R^*})\right] + \beta_\ell^2 \left|A_S\right|^2,$$

$$\begin{split} J_{2s} &= \quad \frac{\beta_{\ell}^2}{4} \left[ |A_{\perp}^L|^2 + |A_{\parallel}^R|^2 + |A_{\perp}^R|^2 + |A_{\parallel}^R|^2 \right], \qquad J_{2c} = -\beta_{\ell}^2 \left[ |A_0^L|^2 + |A_0^R|^2 \right], \\ J_{2s} &= \quad \frac{1}{\beta_{\ell}^2} \left[ |A_{\perp}^L|^2 - |A_{\parallel}^L|^2 + |A_{\parallel}^R|^2 - |A_{\parallel}^R|^2 \right], \qquad J_{4} = \frac{1}{-\beta_{\ell}^2} \left[ \operatorname{Re}(A_0^L A_{\parallel}^{L*} + A_0^R A_{\parallel}^{R*}) \right], \end{split}$$

$$J_5 \quad = \quad \sqrt{2}\beta_\ell \, \left[ {\rm Re}(A_0^L A_\perp^{L\,*} - A_0^R A_\perp^{R\,*}) - \frac{m_\ell}{\sqrt{q^2}} \, {\rm Re}(A_\parallel^L A_S^* + A_\parallel^{R\,*} A_S) \right],$$

$$J_{6s} = 2\beta_{\ell} \left[ \operatorname{Re}(A_{\parallel}^{L}A_{\perp}^{L*} - A_{\parallel}^{R}A_{\perp}^{R*}) \right], \qquad \qquad J_{6c} = 4\beta_{\ell} \frac{m_{\ell}}{\sqrt{q^{2}}} \operatorname{Re}(A_{0}^{L}A_{S}^{*} + A_{0}^{R*}A_{S})$$

$$J_7 \quad = \quad \sqrt{2}\beta_\ell \left[ \mathrm{Im}(\mathbf{A}_0^{\mathrm{L}}\mathbf{A}_\parallel^{\mathrm{L}\,*} - \mathbf{A}_0^{\mathrm{R}}\mathbf{A}_\parallel^{\mathrm{R}\,*}) + \frac{\mathbf{m}_\ell}{\sqrt{\mathbf{q}^2}} \,\mathrm{Im}(\mathbf{A}_\perp^{\mathrm{L}}\mathbf{A}_{\mathrm{S}}^* - \mathbf{A}_\perp^{\mathrm{R}\,*}\mathbf{A}_{\mathrm{S}})) \right],$$

$$J_8 = \frac{1}{\sqrt{2}} \beta_\ell^2 \left[ \operatorname{Im}(\mathbf{A}_0^{\mathbf{L}} \mathbf{A}_\perp^{\mathbf{L}\;*} + \mathbf{A}_0^{\mathbf{R}} \mathbf{A}_\perp^{\mathbf{R}\;*}) \right], \qquad \qquad J_9 = \beta_\ell^2 \left[ \operatorname{Im}(\mathbf{A}_{\parallel}^{\mathbf{L}\;*} \mathbf{A}_\perp^{\mathbf{L}} + \mathbf{A}_{\parallel}^{\mathbf{R}\;*} \mathbf{A}_\perp^{\mathbf{R}}) \right]$$

#### Link to effective operators

 $\Rightarrow$  So here is where the magic happens. At leading order the amplitudes can be written as (soft form factors):

$$A_{\perp}^{L,R} = \sqrt{2}Nm_{B}(1-\hat{s}) \bigg[ (\mathcal{C}_{9}^{\rm eff} + \mathcal{C}_{9}^{\rm eff'}) \mp (\mathcal{C}_{10} + \mathcal{C}_{10}') + \frac{2\hat{m}_{b}}{\hat{s}} (\mathcal{C}_{7}^{\rm eff} + \mathcal{C}_{7}^{\rm eff'}) \bigg] \xi_{\perp}(E_{K}^{*})$$

$$A_{\parallel}^{L,R} \quad = \quad -\sqrt{2}Nm_B(1-\hat{s})\left[ (\mathcal{C}_9^{\mathrm{eff}} - \mathcal{C}_9^{\mathrm{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}}(\mathcal{C}_7^{\mathrm{eff}} - \mathcal{C}_7^{\mathrm{eff}}) \right] \xi_{\perp}(E_K^*)$$

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where  $\hat{s} = q^2/m_B^2$ ,  $\hat{m}_i = m_i/m_B$ . The  $\xi_{\parallel,\perp}$  are the form factors.

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where  $\hat{s} = q^2/m_B^2$ ,  $\hat{m}_i = m_i/m_B$ . The  $\xi_{\parallel,\perp}$  are the form factors.  $\Rightarrow$  Now we can construct observables that cancel the  $\xi$  form factors at leading order:

$$P_5' = \frac{J_5 + J_5}{2\sqrt{-(J_2^c + \bar{J}_2^c)(J_2^s + \bar{J}_2^s)}}$$

## $B^0 \rightarrow K^* \mu^- \mu^+$ kinematics

⇒ The kinematics of  $B^0 \to K^* \mu^- \mu^+$  decay is described by three angles  $\theta_l$ ,  $\theta_k$ ,  $\phi$  and invariant mass of the dimuon system  $(q^2)$ .

⇒  $\cos \theta_k$ : the angle between the direction of the kaon in the  $K^*$  ( $\overline{K}^*$ ) rest frame and the direction of the  $K^*$  ( $\overline{K}^*$ ) in the  $B^0$  ( $\overline{B}^0$ ) rest frame. ⇒  $\cos \theta_l$ : the angle between the direction of the  $\mu^-$  ( $\mu^+$ ) in the dimuon rest frame and the direction of the dimuon in the  $B^0$  ( $\overline{B}^0$ ) rest frame.

⇒  $\phi$ : the angle between the plane containing the  $\mu^-$  and  $\mu^+$  and the plane containing the kaon and pion from the  $K^*$ .



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$$\begin{split} \frac{d^4 \Gamma}{dq^2 \operatorname{dcos} \theta_K \operatorname{dcos} \theta_l d\phi} &= \frac{9}{32\pi} \left[ J_{1s} \sin^2 \theta_K + J_{1c} \cos^2 \theta_K + (J_{2s} \sin^2 \theta_K + J_{2c} \cos^2 \theta_K) \cos 2\theta_l \right. \\ &+ J_3 \sin^2 \theta_K \sin^2 \theta_l \cos 2\phi + J_4 \sin 2\theta_K \sin 2\theta_l \cos \phi + J_5 \sin 2\theta_K \sin \theta_l \cos \phi \\ &+ (J_{6s} \sin^2 \theta_K + J_{6c} \cos^2 \theta_K) \cos \theta_l + J_7 \sin 2\theta_K \sin \theta_l \sin \phi + J_8 \sin 2\theta_K \sin 2\theta_l \sin \phi \\ &+ J_9 \sin^2 \theta_K \sin^2 \theta_l \sin 2\phi \right], \end{split}$$

 $\Rightarrow$  This is the most general expression of this kind of decay.  $\Rightarrow$  The *CP* averaged angular observables are defined:

$$S_i = \frac{J_i + \bar{J}_i}{(d\Gamma + d\bar{\Gamma})/dq^2}$$

#### Link to effective operators

 $\Rightarrow \text{The observables } J_i \text{ are bilinear combinations of transversity amplitudes: } A^{L,R}_{\perp}, \ A^{L,R}_{\parallel}, \ A^{L,R}_{0}.$ 

 $\Rightarrow$  So here is where the magic happens. At leading order the amplitudes can be written as:

$$\begin{split} A_{\perp}^{L,R} &= -\sqrt{2}Nm_B(1-\hat{s}) \left[ (\mathcal{C}_9^{\rm eff} + \mathcal{C}_9^{\rm eff\prime}) \mp (\mathcal{C}_{10} + \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\rm eff} + \mathcal{C}_7^{\rm eff\prime}) \right] \xi_{\perp}(E_{K^*}) \\ A_{\parallel}^{L,R} &= -\sqrt{2}Nm_B(1-\hat{s}) \left[ (\mathcal{C}_9^{\rm eff} - \mathcal{C}_9^{\rm eff\prime}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + \frac{2\hat{m}_b}{\hat{s}} (\mathcal{C}_7^{\rm eff} - \mathcal{C}_7^{\rm eff\prime}) \right] \xi_{\perp}(E_{K^*}) \end{split}$$

$$A_{0}^{L,R} = -\frac{Nm_{B}(1-\hat{s})^{2}}{2\hat{m}_{K}^{*}\sqrt{\hat{s}}} \bigg[ (\mathcal{C}_{9}^{\mathrm{eff}} - \mathcal{C}_{9}^{\mathrm{eff}}) \mp (\mathcal{C}_{10} - \mathcal{C}_{10}') + 2\hat{m}_{b}(\mathcal{C}_{7}^{\mathrm{eff}} - \mathcal{C}_{7}^{\mathrm{eff}}) \bigg] \xi_{\parallel}(E_{K}^{*}),$$

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#### Symmetries in $B \rightarrow K^* \mu \mu$

 $\Rightarrow$  We have 12 angular coefficients (S<sub>i</sub>).

 $\Rightarrow$  There exist 4 symmetry transformations that leave the angular distributions unchanged:

$$\boldsymbol{n}_{\parallel} = \begin{pmatrix} \boldsymbol{A}_{\parallel}^L \\ \boldsymbol{A}_{\parallel}^{R*} \end{pmatrix}, \quad \boldsymbol{n}_{\perp} = \begin{pmatrix} \boldsymbol{A}_{\perp}^L \\ -\boldsymbol{A}_{\perp}^{R*} \end{pmatrix}, \quad \boldsymbol{n}_0 = \begin{pmatrix} \boldsymbol{A}_0^L \\ \boldsymbol{A}_0^{R*} \end{pmatrix}.$$

$$n_i' = U n_i = \left[ \begin{array}{cc} e^{i\phi_L} & 0 \\ 0 & e^{-i\phi_R} \end{array} \right] \left[ \begin{array}{cc} \cos\theta & -\sin\theta \\ \sin\theta & \cos\theta \end{array} \right] \left[ \begin{array}{cc} \cosh i\tilde{\theta} & -\sinh i\tilde{\theta} \\ -\sinh i\tilde{\theta} & \cosh i\tilde{\theta} \end{array} \right] n_i \, . \label{eq:ni}$$

 $\Rightarrow$  Using this symmetries one can show that there are 8 independent observables. The pdf can be written as:

$$\begin{split} \frac{1}{\mathrm{d}(\Gamma+\bar{\Gamma})/\mathrm{d}q^2} \frac{\mathrm{d}(\Gamma+\bar{\Gamma})}{\mathrm{d}\cos\theta_l \,\mathrm{d}\cos\theta_k \,\mathrm{d}\phi} \bigg|_{\mathrm{P}} &= \frac{9}{32\pi} \left[ \frac{3}{4} (1-F_\mathrm{L}) \sin^2\theta_k \\ &+ F_\mathrm{L} \cos^2\theta_k + \frac{1}{4} (1-F_\mathrm{L}) \sin^2\theta_k \cos 2\theta_l \\ &- F_\mathrm{L} \cos^2\theta_k \cos 2\theta_l + S_3 \sin^2\theta_k \sin^2\theta_l \cos 2\phi \\ &+ S_4 \sin 2\theta_k \sin 2\theta_l \cos \phi + S_5 \sin 2\theta_k \sin \theta_l \cos \phi \\ &+ \frac{4}{3} A_\mathrm{FB} \sin^2\theta_k \cos \theta_l + S_7 \sin 2\theta_k \sin \theta_l \sin \phi \\ &+ S_8 \sin 2\theta_k \sin 2\theta_l \sin \phi + S_9 \sin^2\theta_k \sin^2\theta_l \sin 2\phi \right]. \end{split}$$